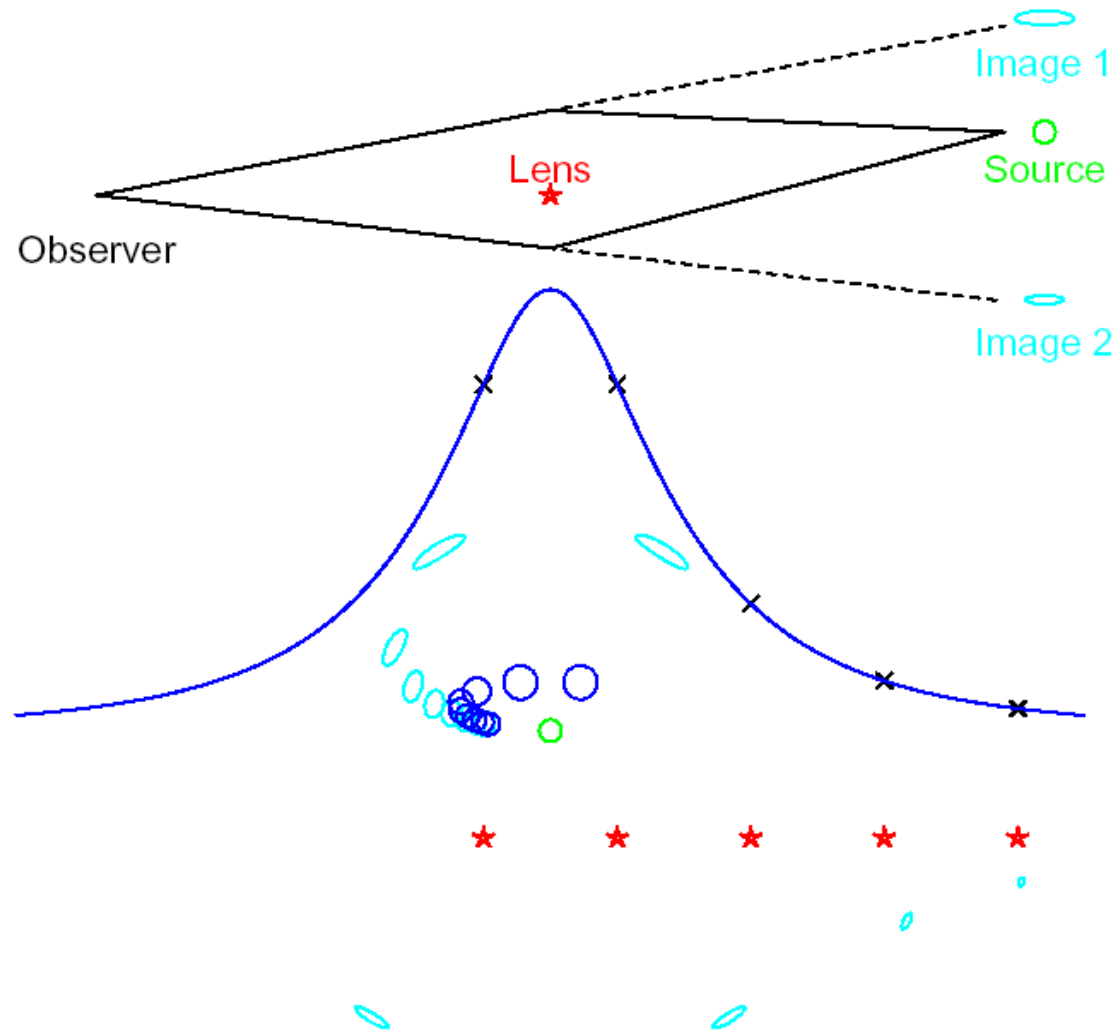


# $\mu$ FUN Planet Detections

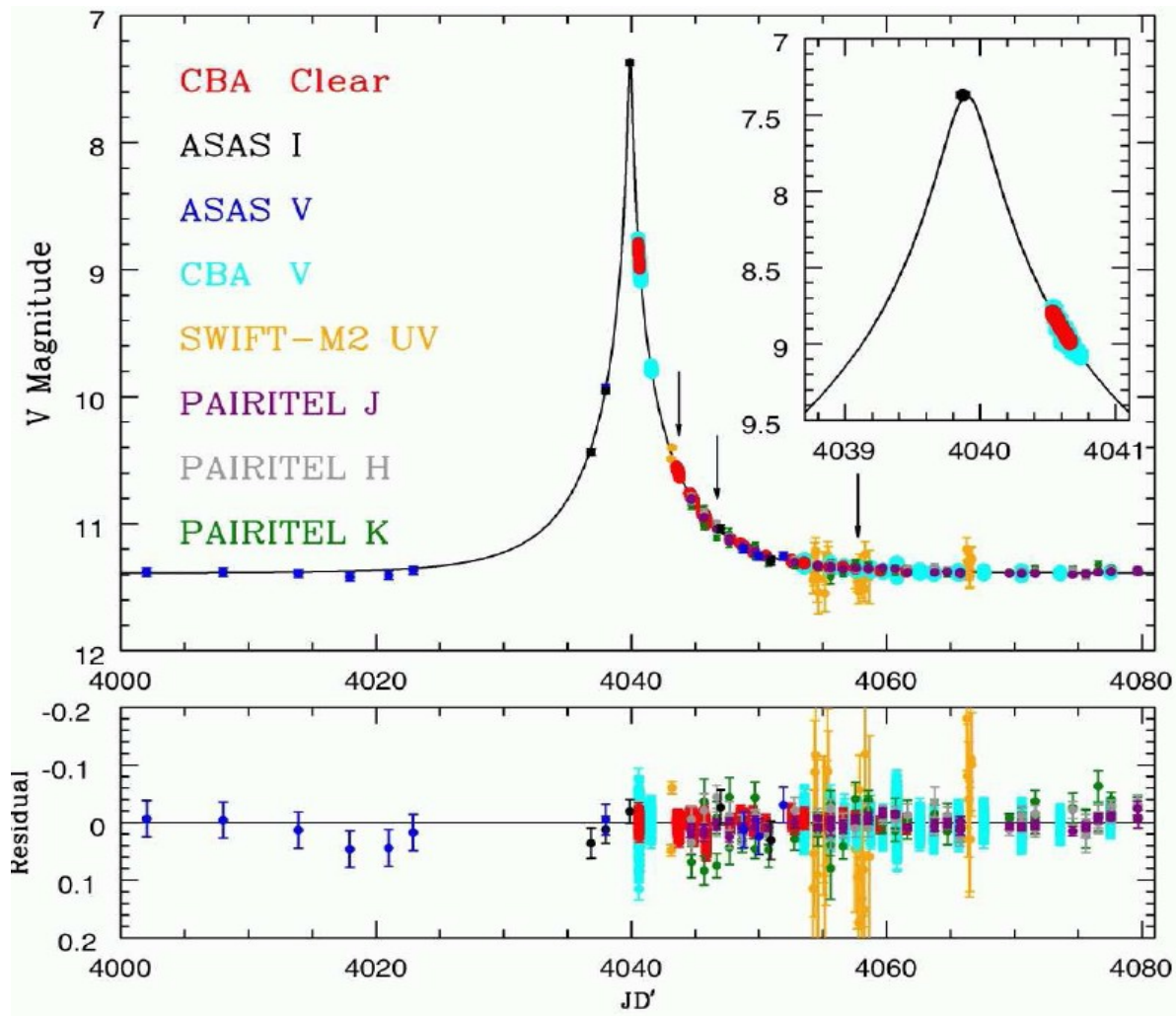
2005-2008

Andy Gould (Ohio State)



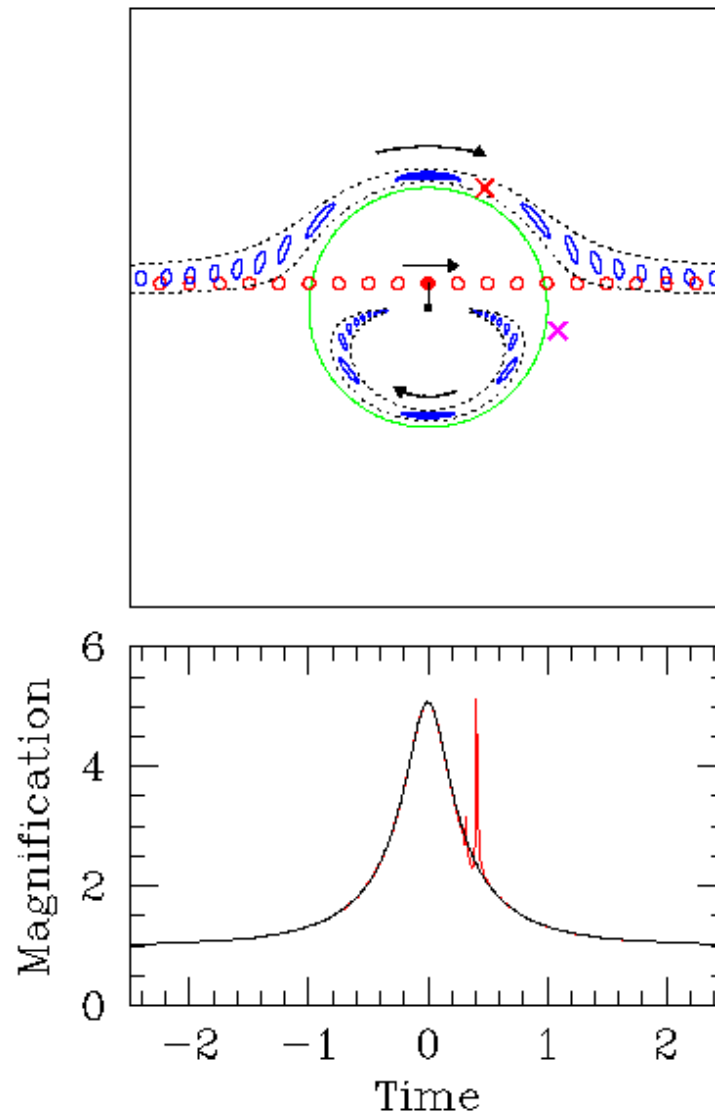
# A0 star GSC 3656-1328

A=40, V=7 at 1 kpc



Gaudi et al. 2008, ApJ

# How Microlensing Finds Planets



# Mao & Paczynski

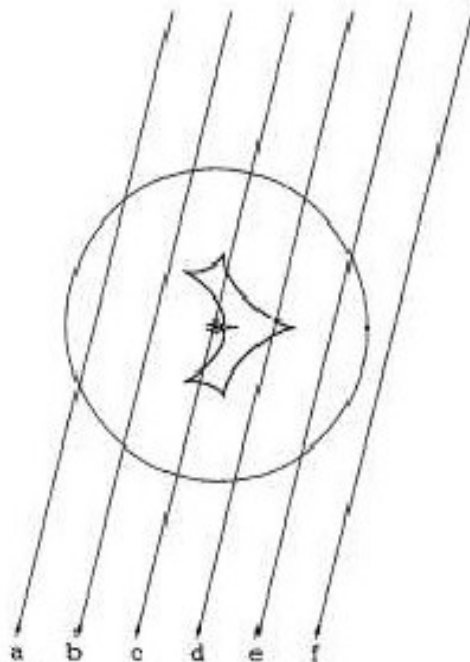
## Microlens Planet Searches

### GRAVITATIONAL MICROLENSING BY DOUBLE STARS AND PLANETARY SYSTEMS

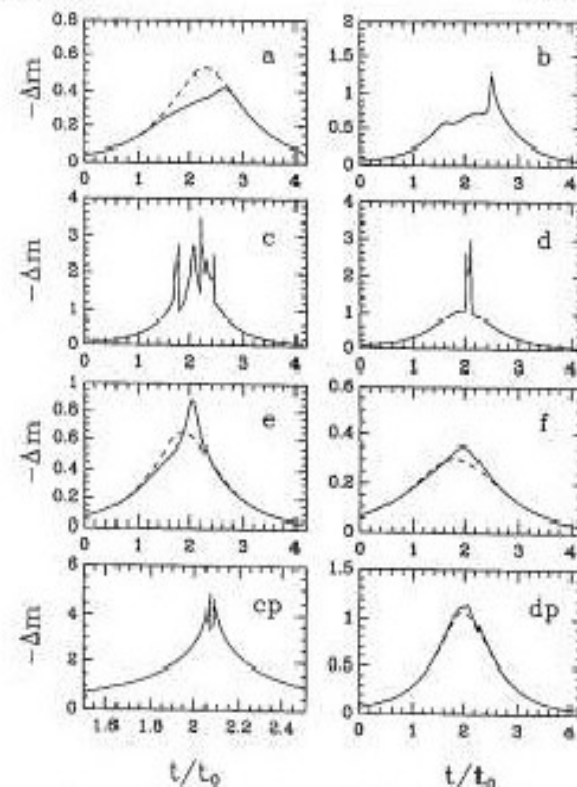
SHUDE MAO AND BOHDAN PACZYŃSKI

Princeton University Observatory, Princeton, NJ 08544

Received 1991 March 12; accepted 1991 April 2



1.—Geometry of microlensing by a binary, as seen in the sky. The  $y$  star of  $1 M_{\odot}$  is located at the center of the figure, and the secondary of  $y$  or  $0.001 M_{\odot}$  is located on the right, on the Einstein ring of the  $y$ . The radius of the ring is 1.0 mas for a source located at a distance of 8 kpc. The two complicated shapes around the primary are



the lens. The effect is strong even if the companion is a planet. A massive search for microlensing of the Galactic bulge stars may lead to a discovery of the first extrasolar planetary systems.

# Survey + Follow-Up

## DISCOVERING PLANETARY SYSTEMS THROUGH GRAVITATIONAL MICROLENSSES

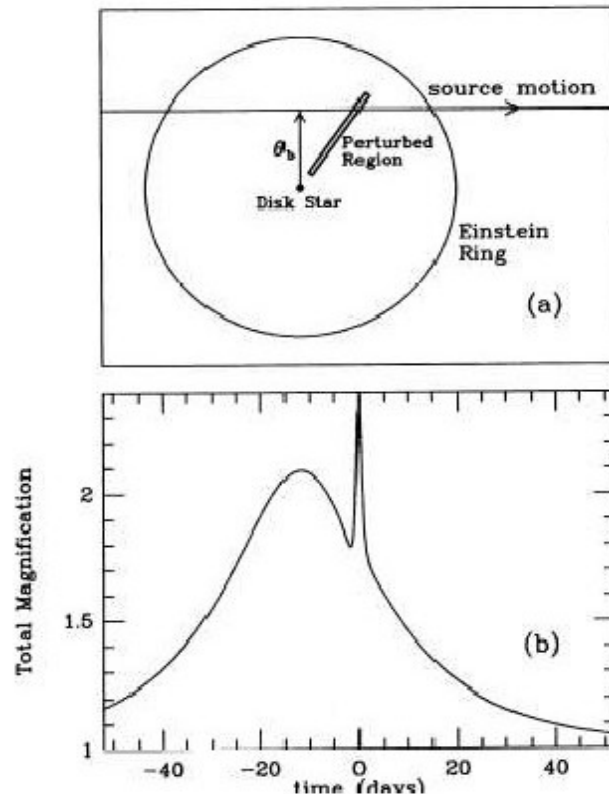
ANDREW GOULD AND ABRAHAM LOEB  
Institute for Advanced Study, Princeton, NJ 08540  
*Received 1991 December 26; accepted 1992 March 9*

### 5. OBSERVATIONAL REQUIREMENTS

Two distinct steps are required to observe a planetary system by microlensing. First, one must single out a disk star which happens to be microlensing a bulge star. Second, one must observe this star often enough to catch the deviation in the light curve due to the planet. The first step involves the observation of millions of bulge stars on the order of once per day. The second step involves the observation of a handful of stars many times per day. In the following we give a rough outline of what is required for each of these steps.

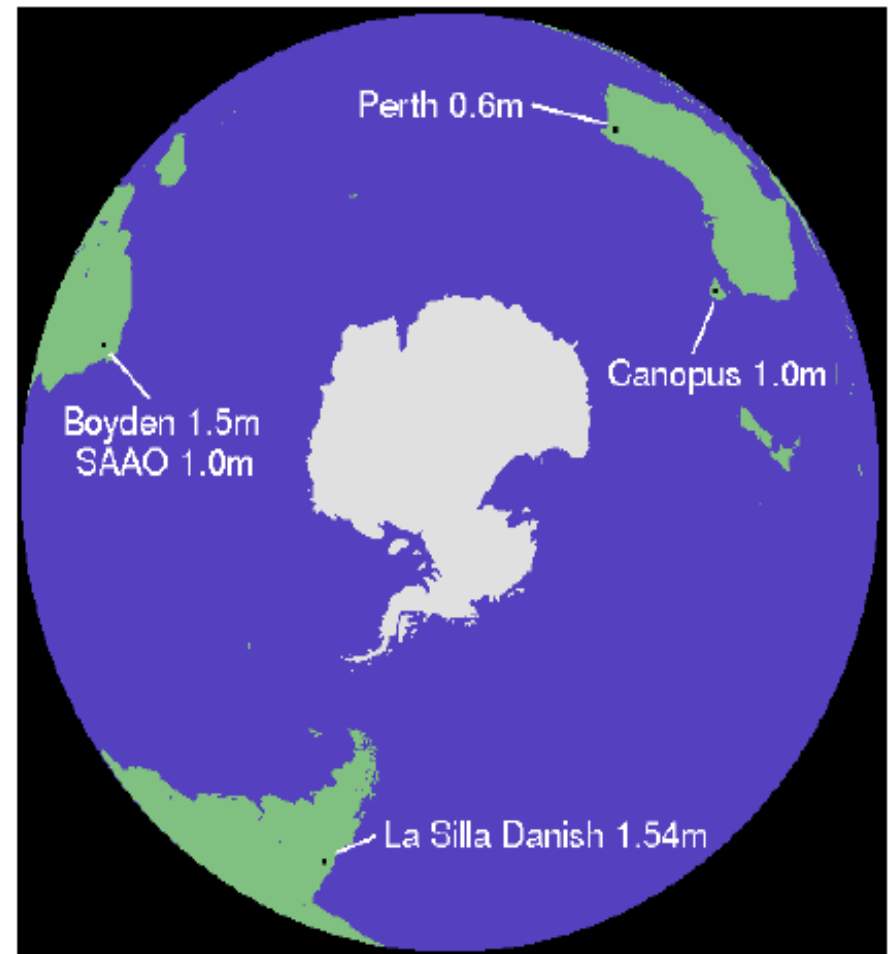
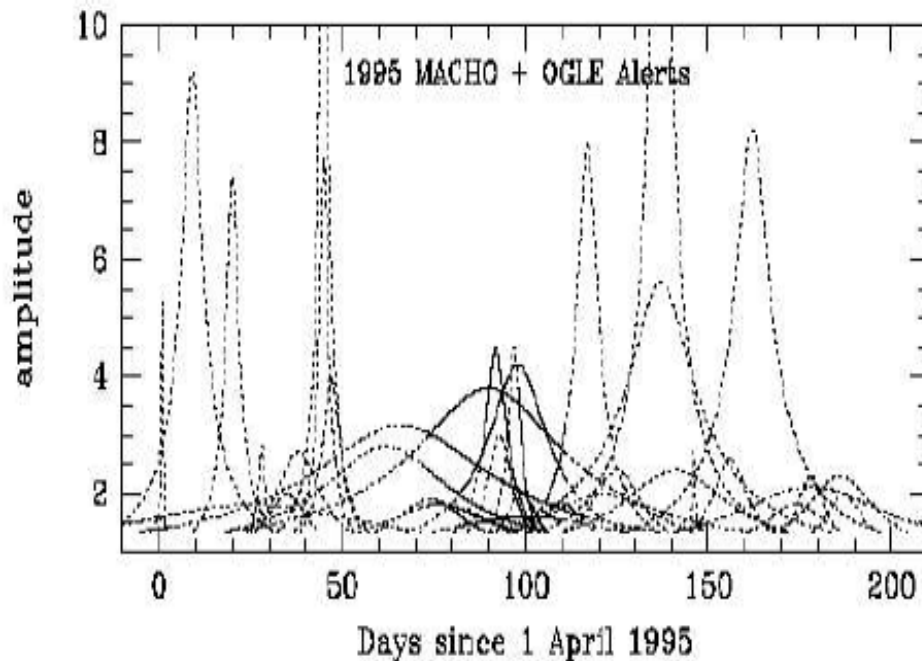
While observations from one site would be useful, there are advantages to be gained by observing from several sites. First,

two telescopes that were totally committed. Third, in view of the fleeting nature of the events, it would seem prudent to build in some redundancy in case of bad weather at a particular site. Thus, the optimal scheme would employ, say, a dozen telescopes. Each of these would be committed to carry out two observations per night. During the near-December season,



# 1995 PLANET Pilot Season

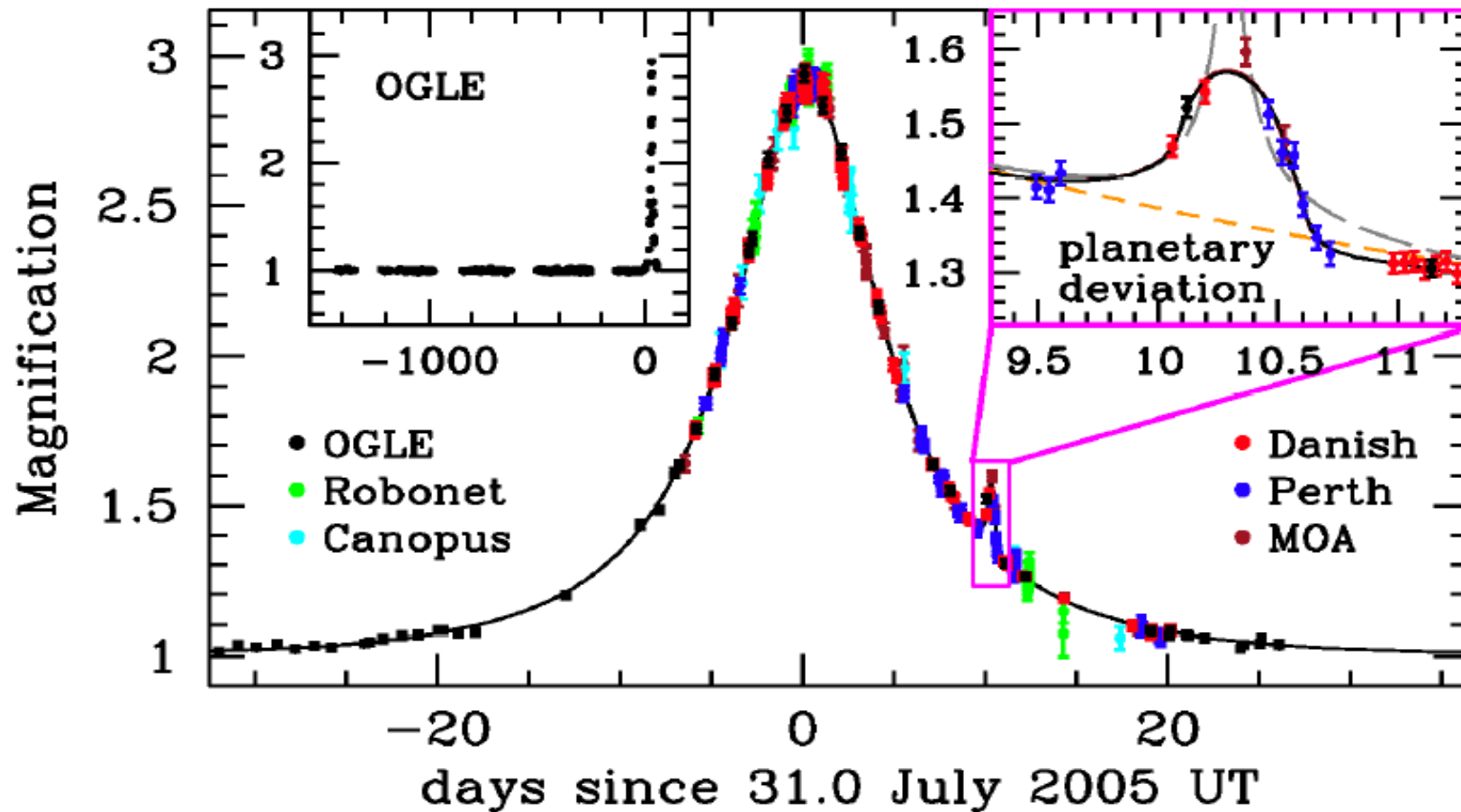
- Albrow et al. 1998  
*ApJ*, 509, 687





# OGLE-2005-BLG-390

## “Classical-Followup” Planetary Caustic

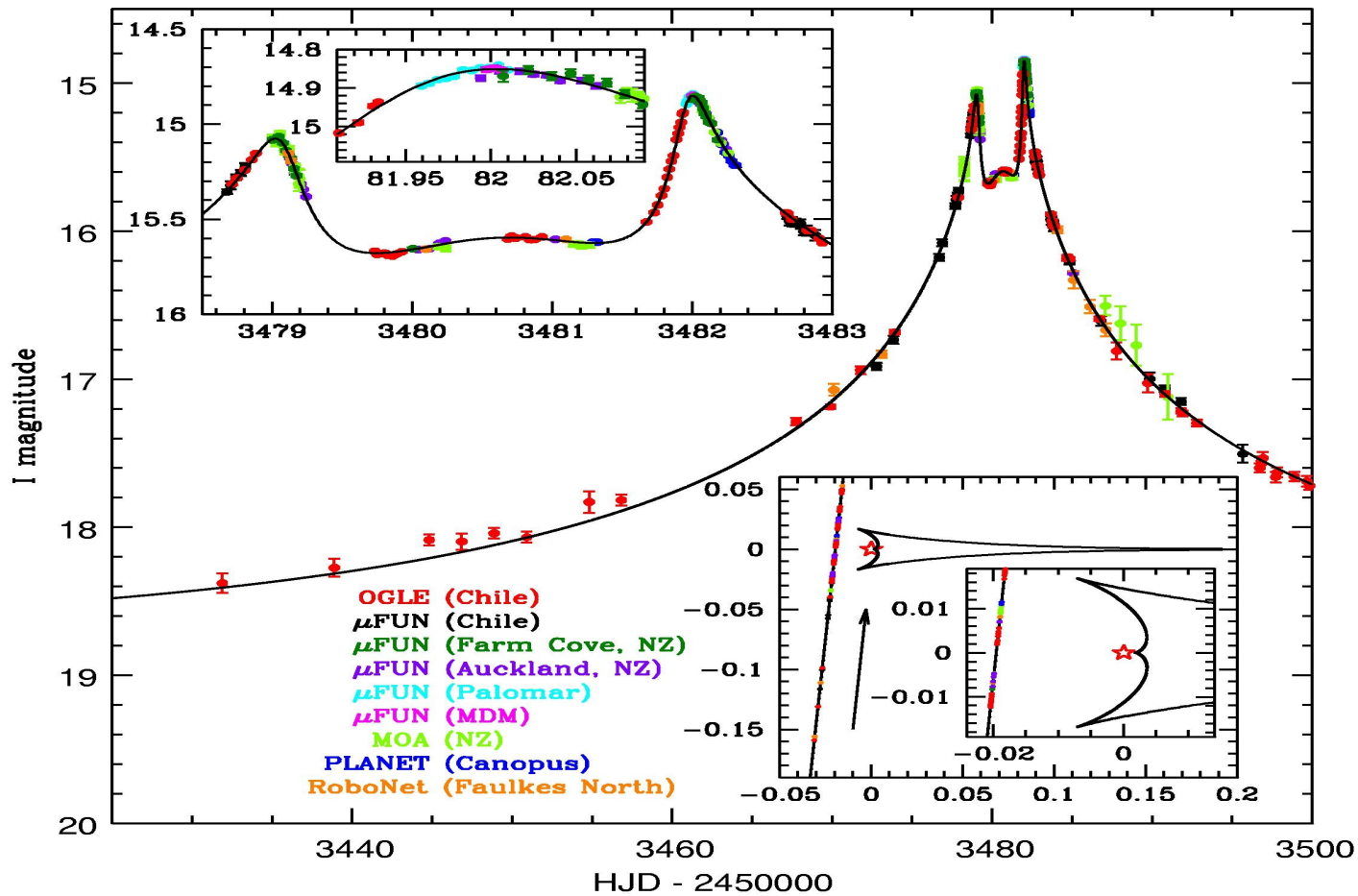


Beaulieu et al. 2006, Nature, 439, 437

# OGLE-2005-BLG-071

## 1<sup>st</sup> $\mu$ FUN Planet:

## Super-Jupiter around M dwarf



Udalski et al. 2005, ApJ, 628, L109



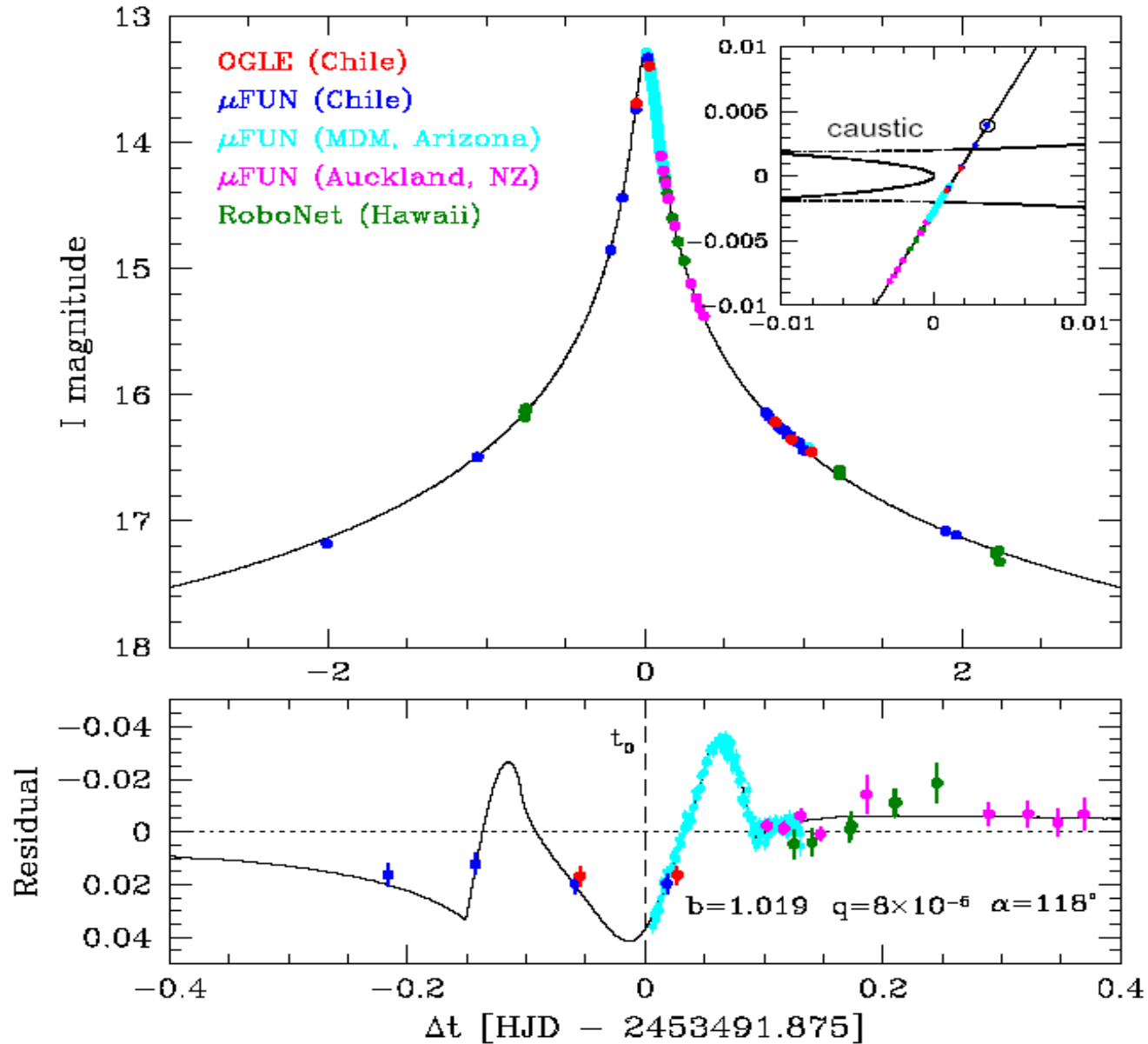
# Amateurs + Professionals

Grant, Ian, Jennie, Phil



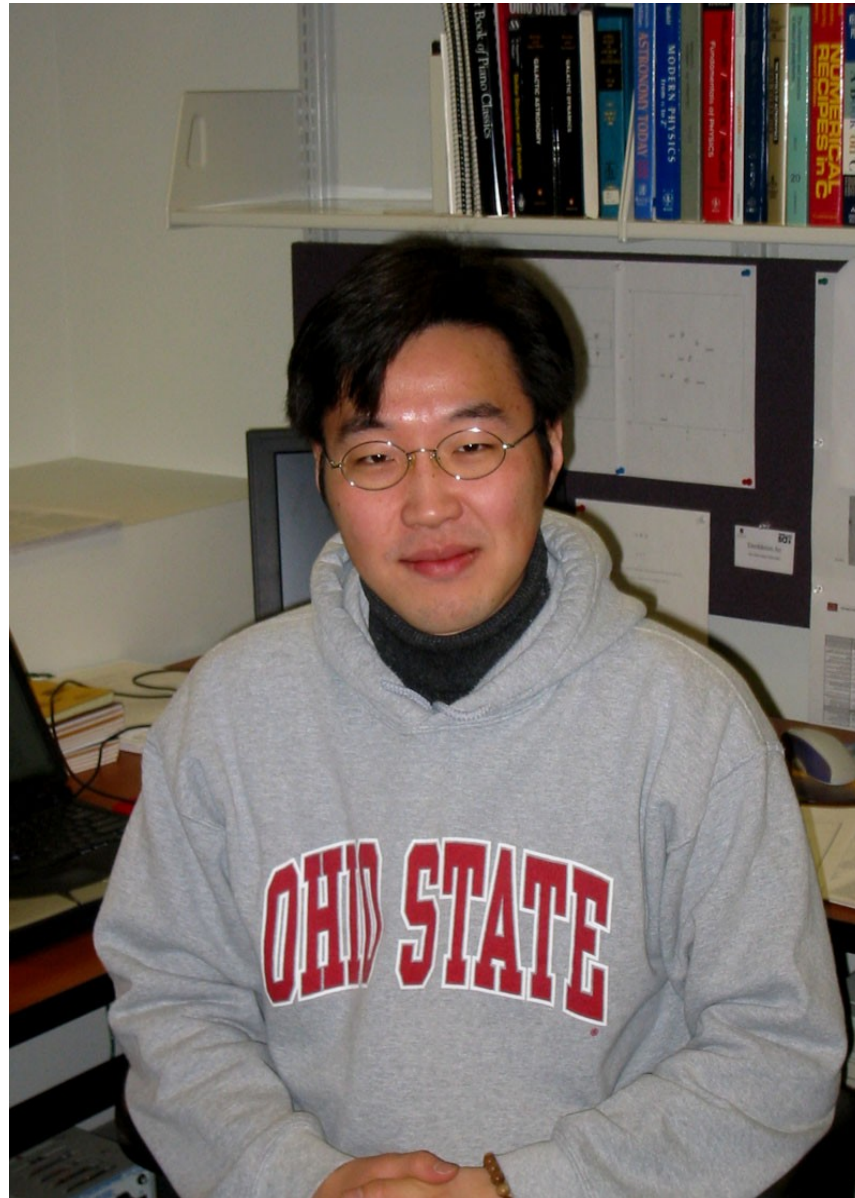
# OGLE-2005-BLG-169:

## Second $\mu$ FUN Planet: Cold Neptune





# Deokkeun An



# Tale of Two Planets

## OGLE-2005-BLG-390

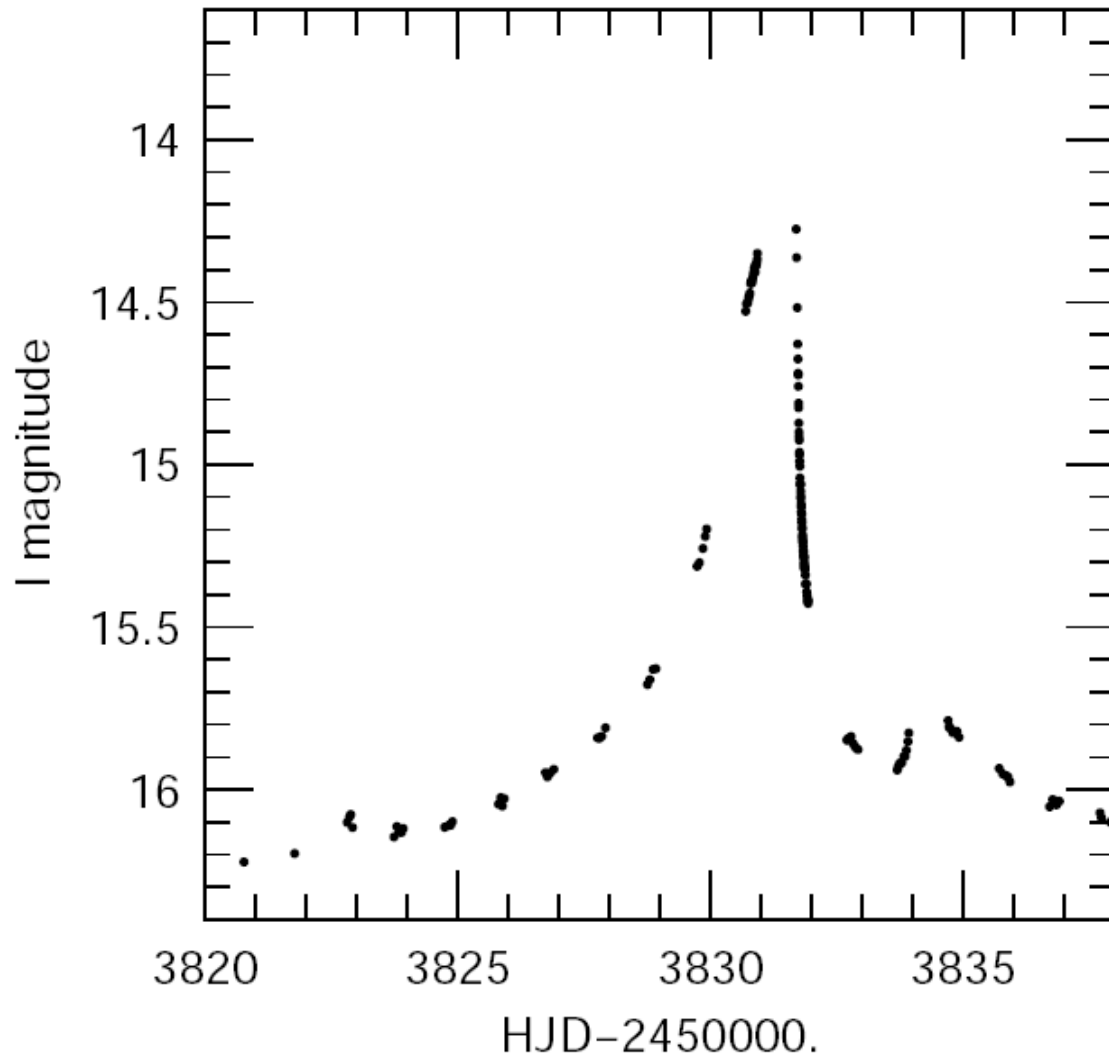
- $q = 8e-5$
- $M_* = 0.2 M_{\text{sun}}$
- $M_p = 5.5 M_{\text{earth}}$
- $D = 7 \text{ kpc}$
- $a = 3 \text{ AU}$
- $T = 50 \text{ K}$
- Low-mag Event
- (1 det)/(4.4 probed)

## OGLE-2005-BLG-169

- $q = 8e-5$
- $M_* = 0.5 M_{\text{sun}}$
- $M_p = 13 M_{\text{earth}}$
- $D = 3 \text{ kpc}$
- $a = 4 \text{ AU}$
- $T = 70 \text{ K}$
- High-mag Event
- (1 det)/(2.25 probed)

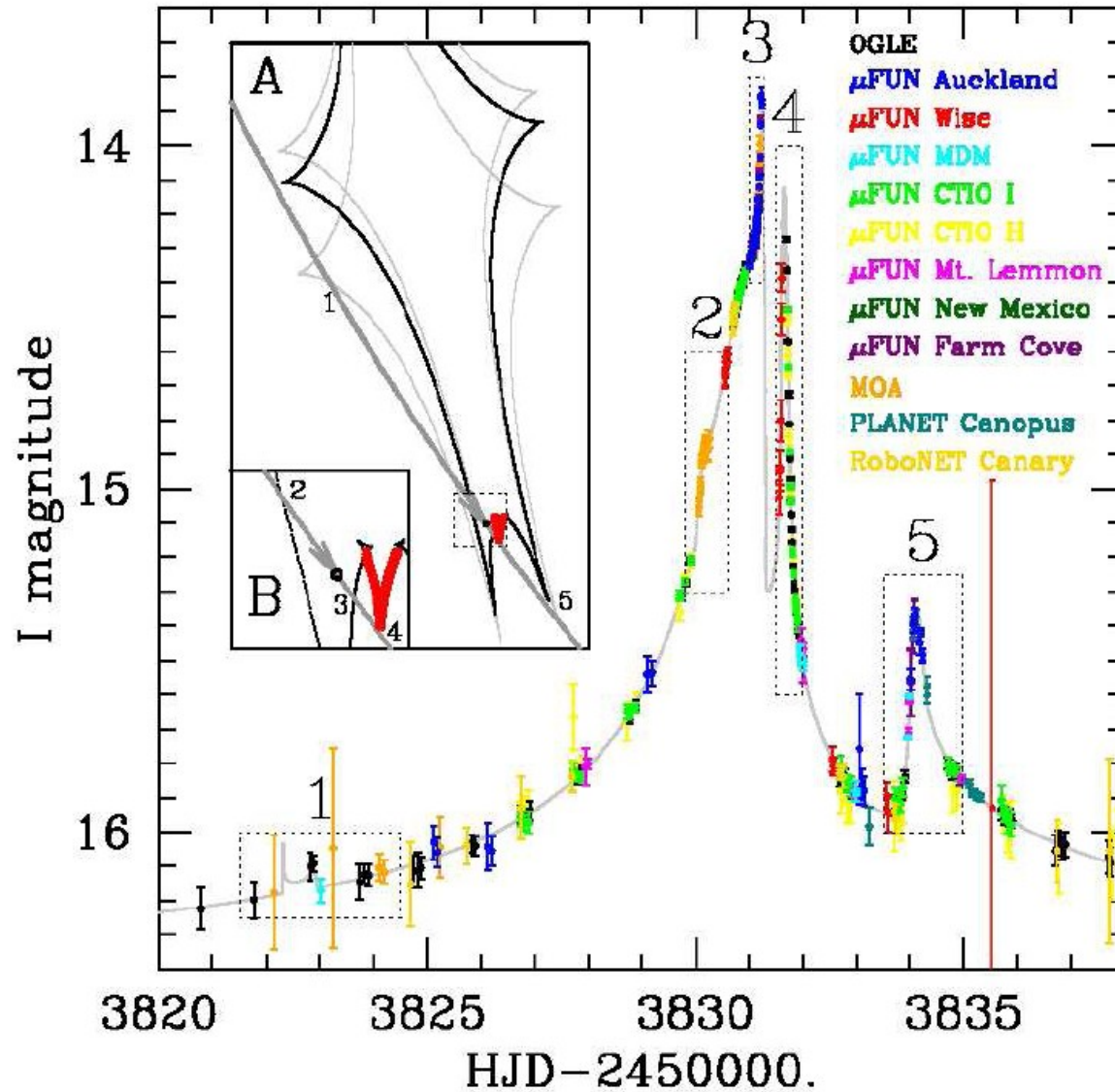
# OGLE-2006-BLG-109:

## Without Followup Observations



# OGLE-2006-BLG-109

## $\mu$ FUN Planets 3+4: Jupiter+Saturn System

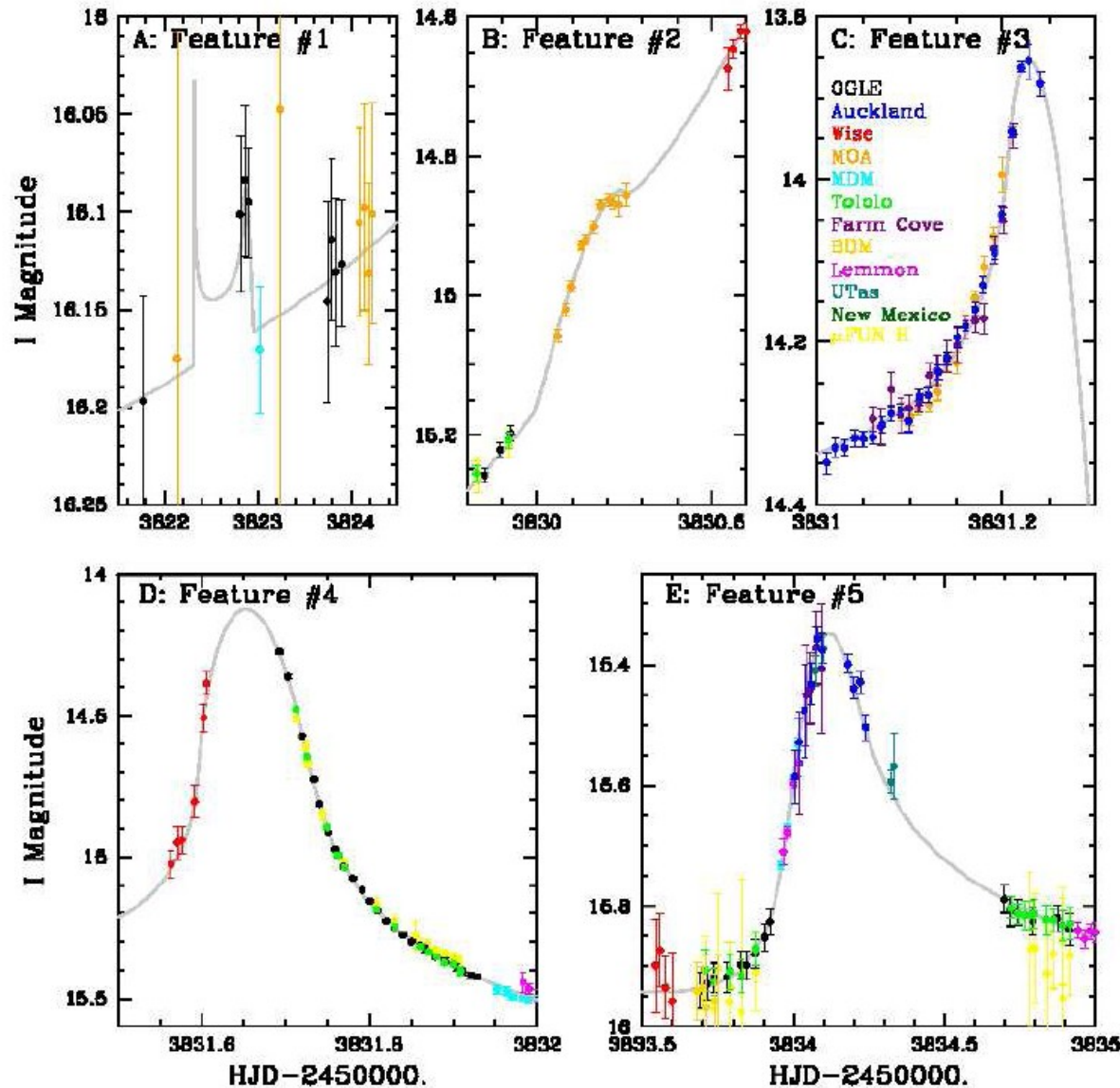


Gaudi et al. 2008, Science, 319, 927



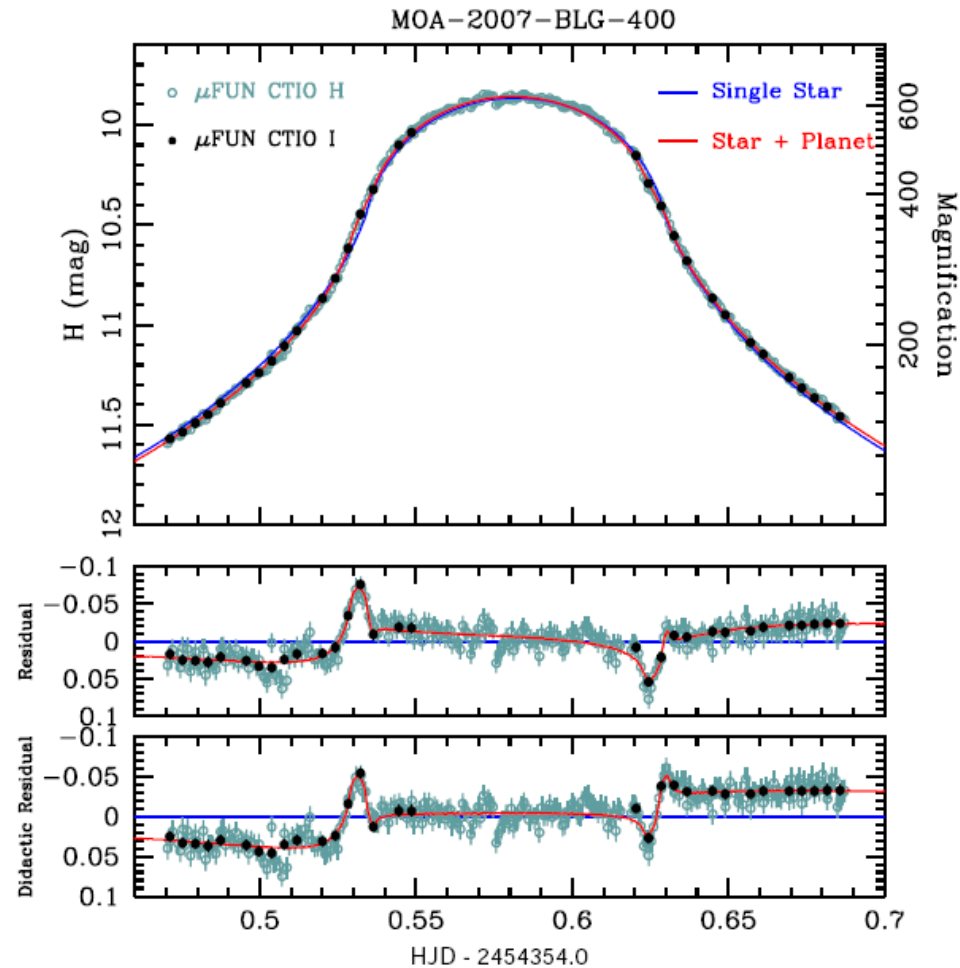
# Five Lightcurve Features

1+2+3+5=Saturn    4=Jupiter

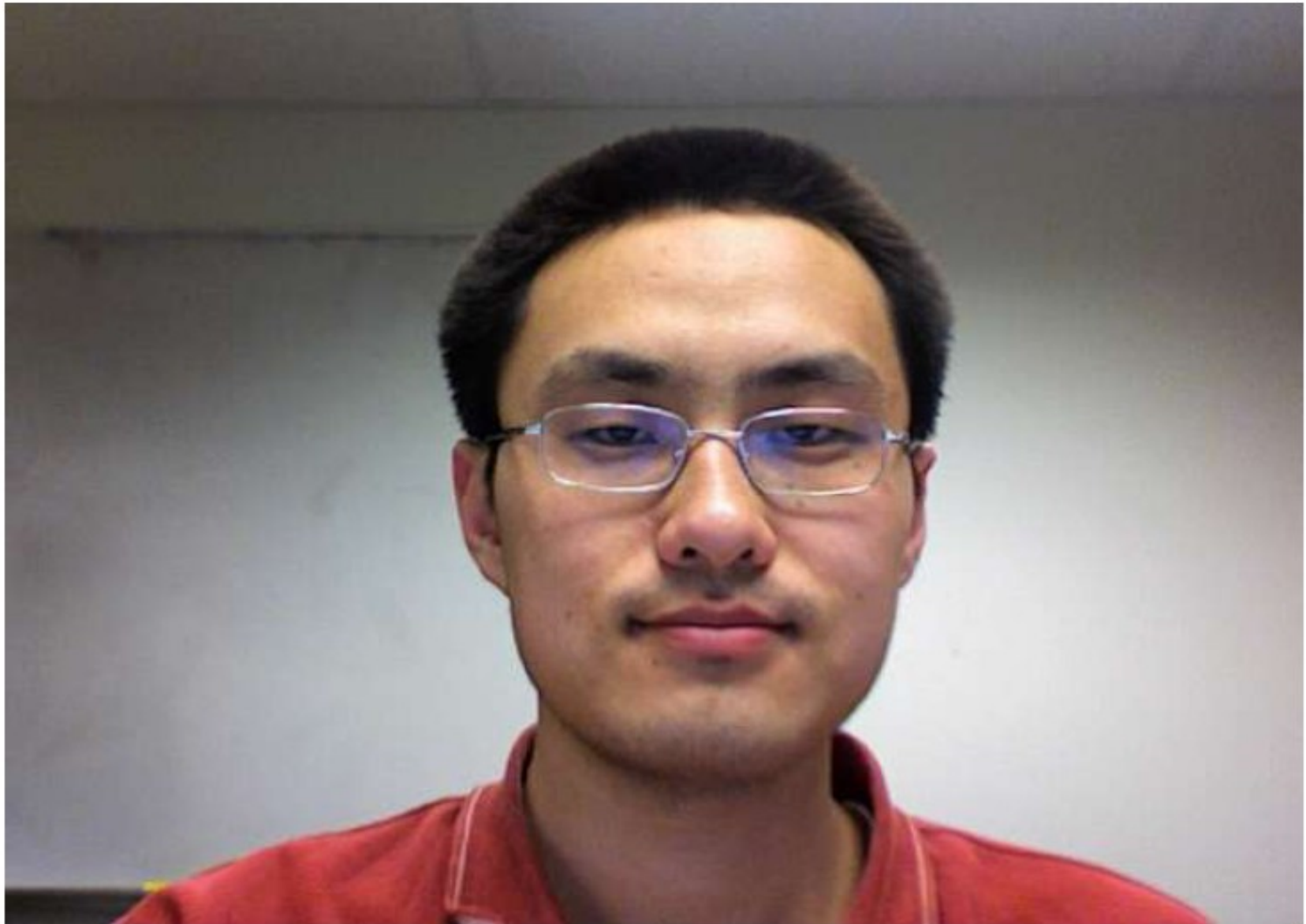


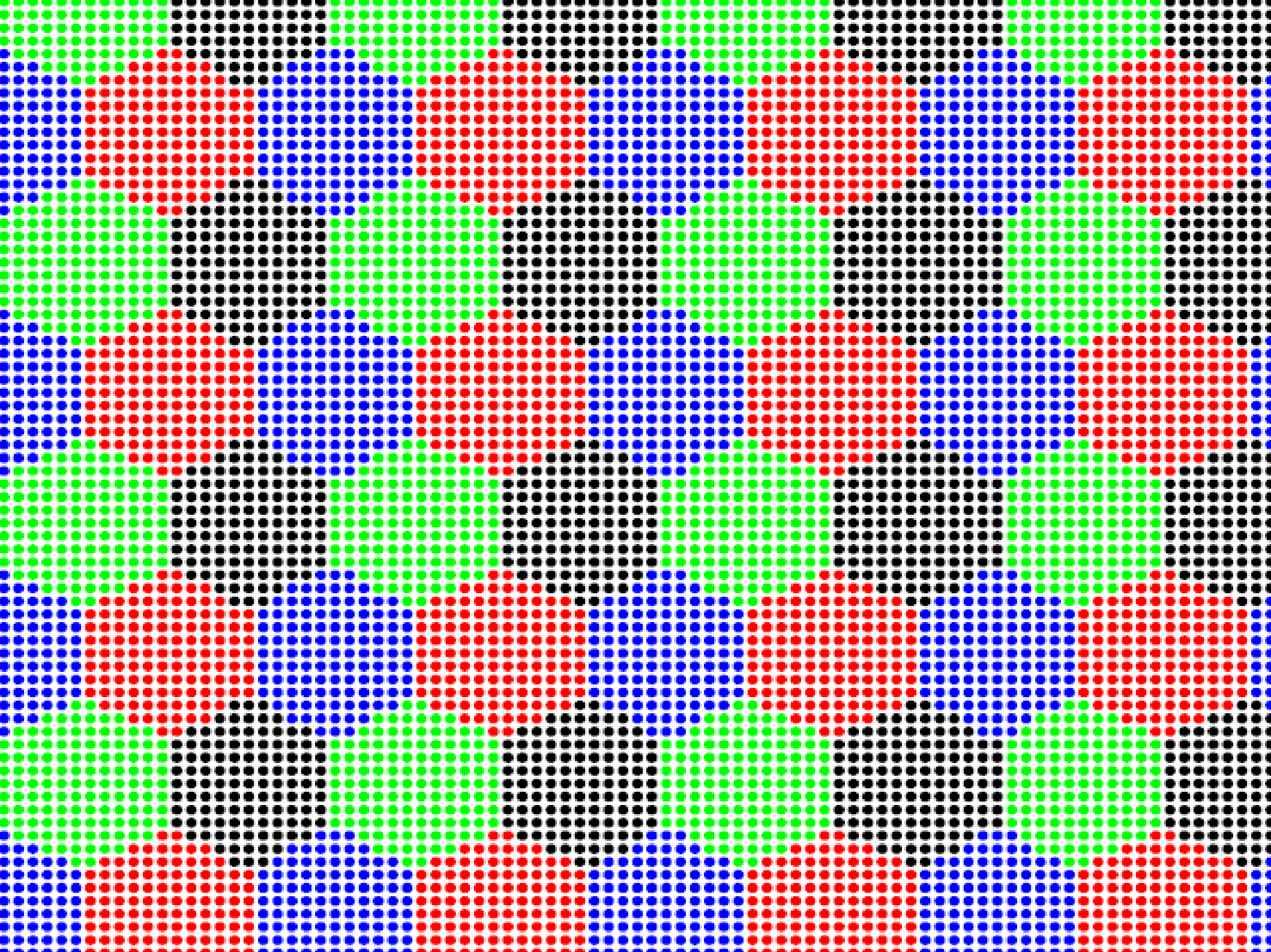
# MOA-2007-BLG-400

## Fifth $\mu$ FUN Planet: “Buried” Jupiter

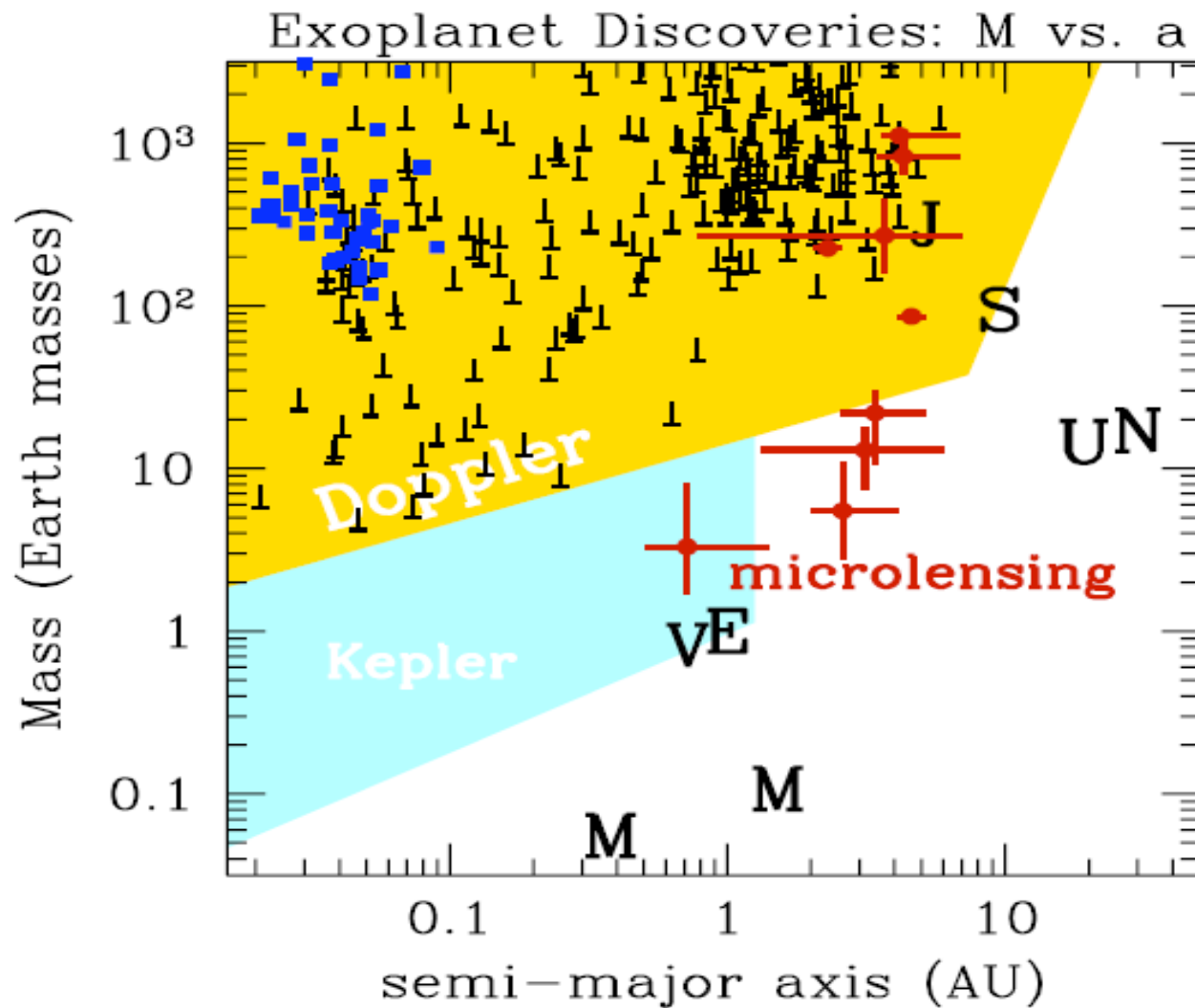


# Subo Dong

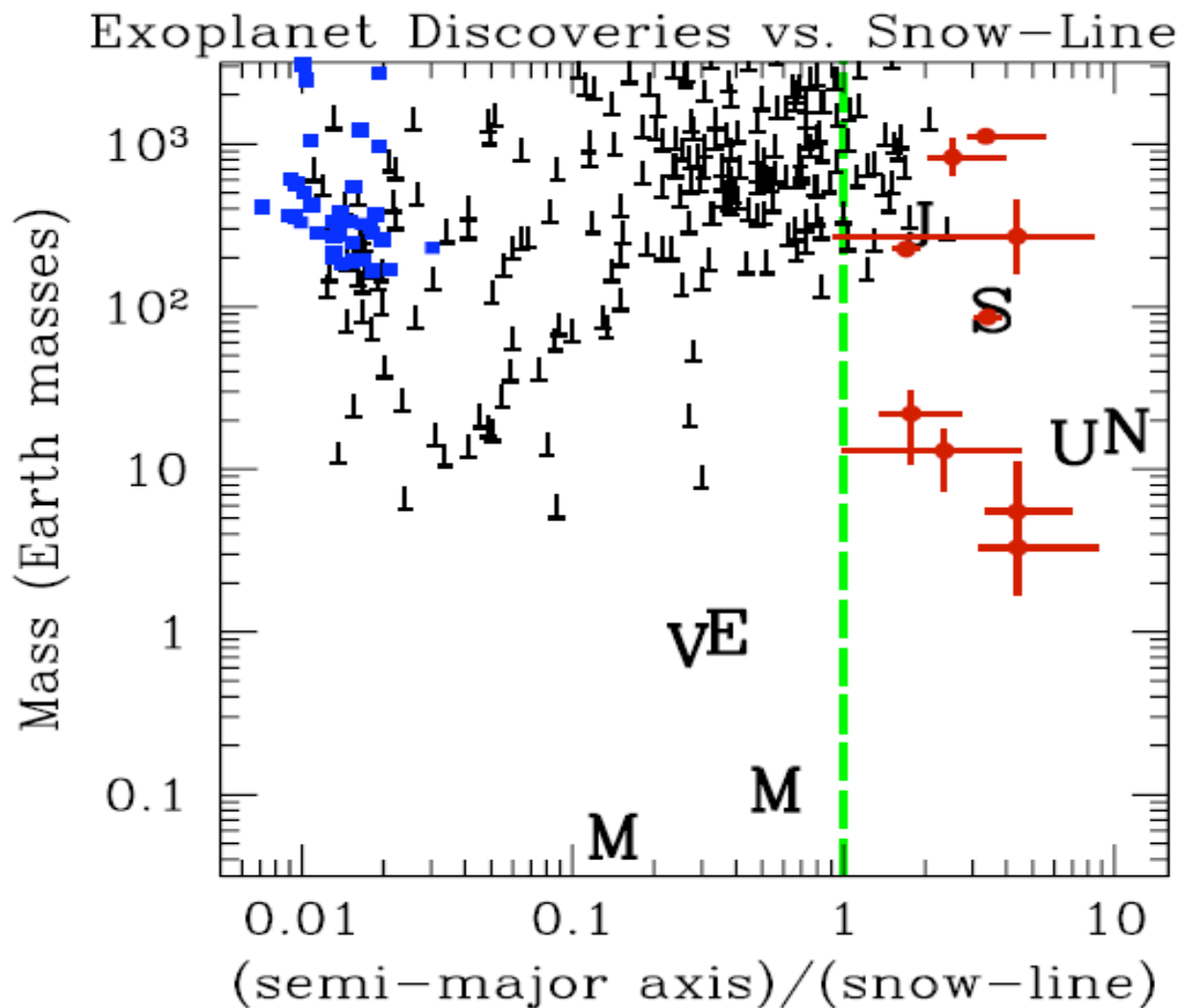




# Microlensing vs. Other Methods

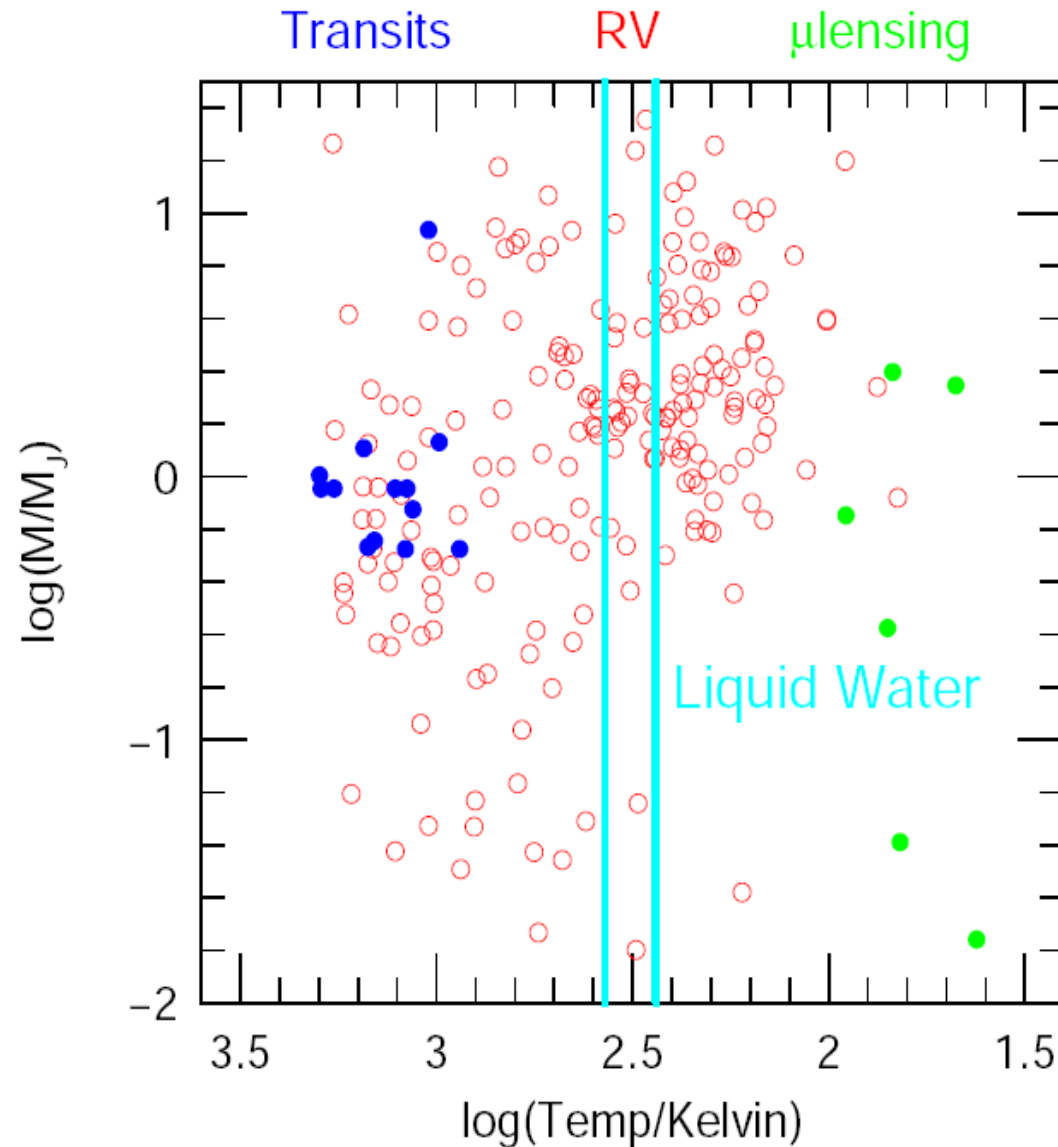


# Microlensing and the “Snow Line”

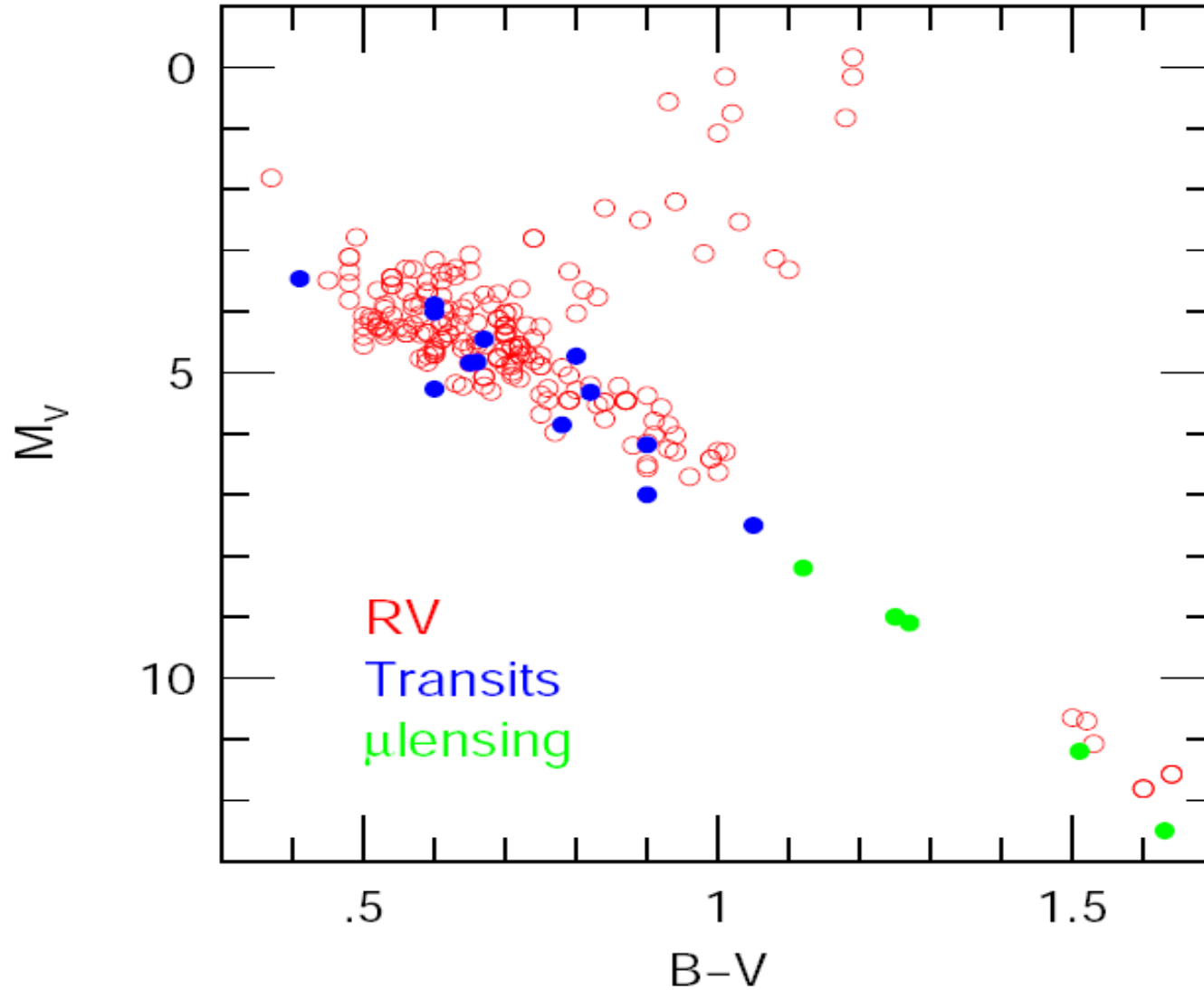




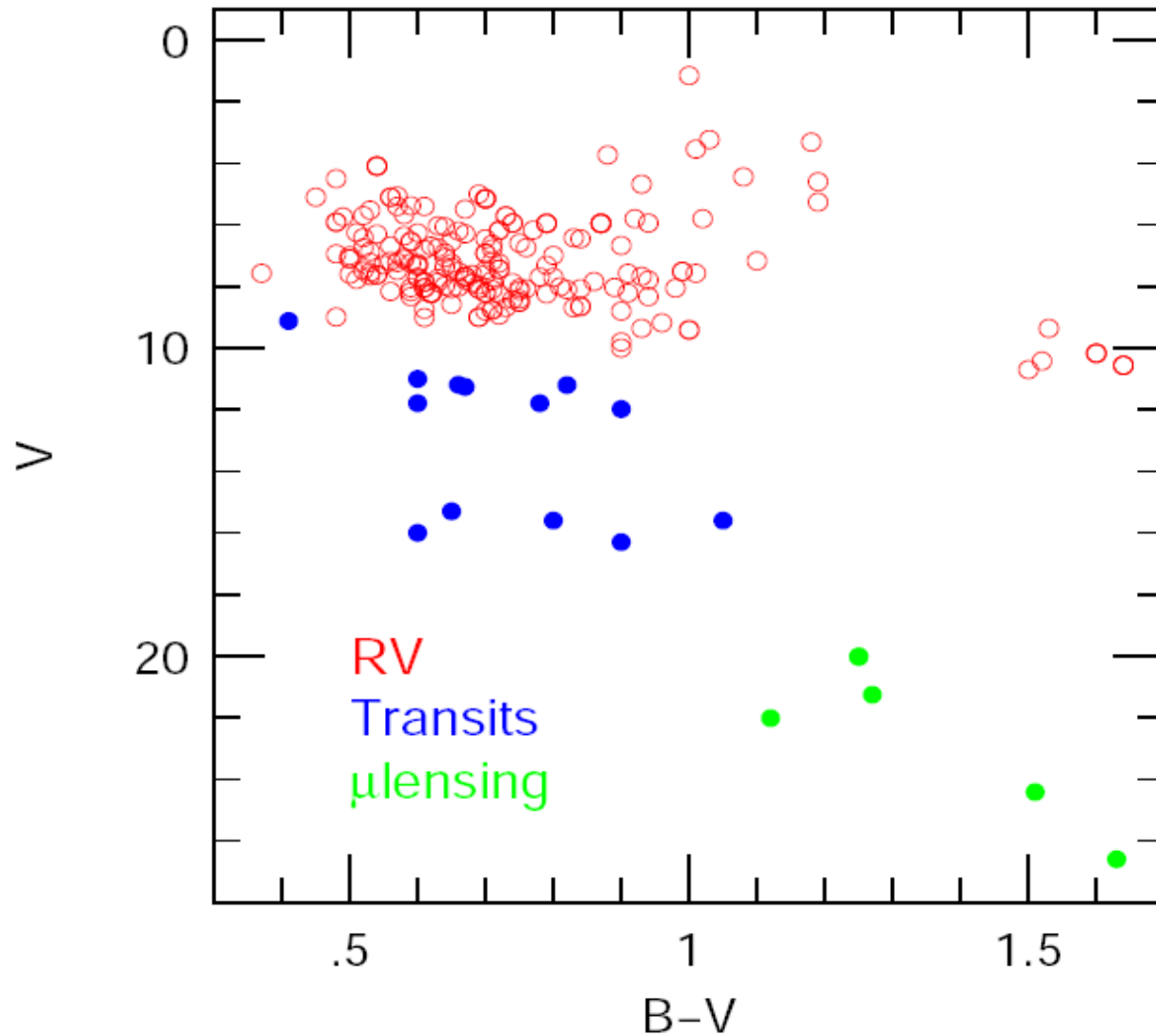
# Equilibrium Temperature



# CMD (Absolute mags)

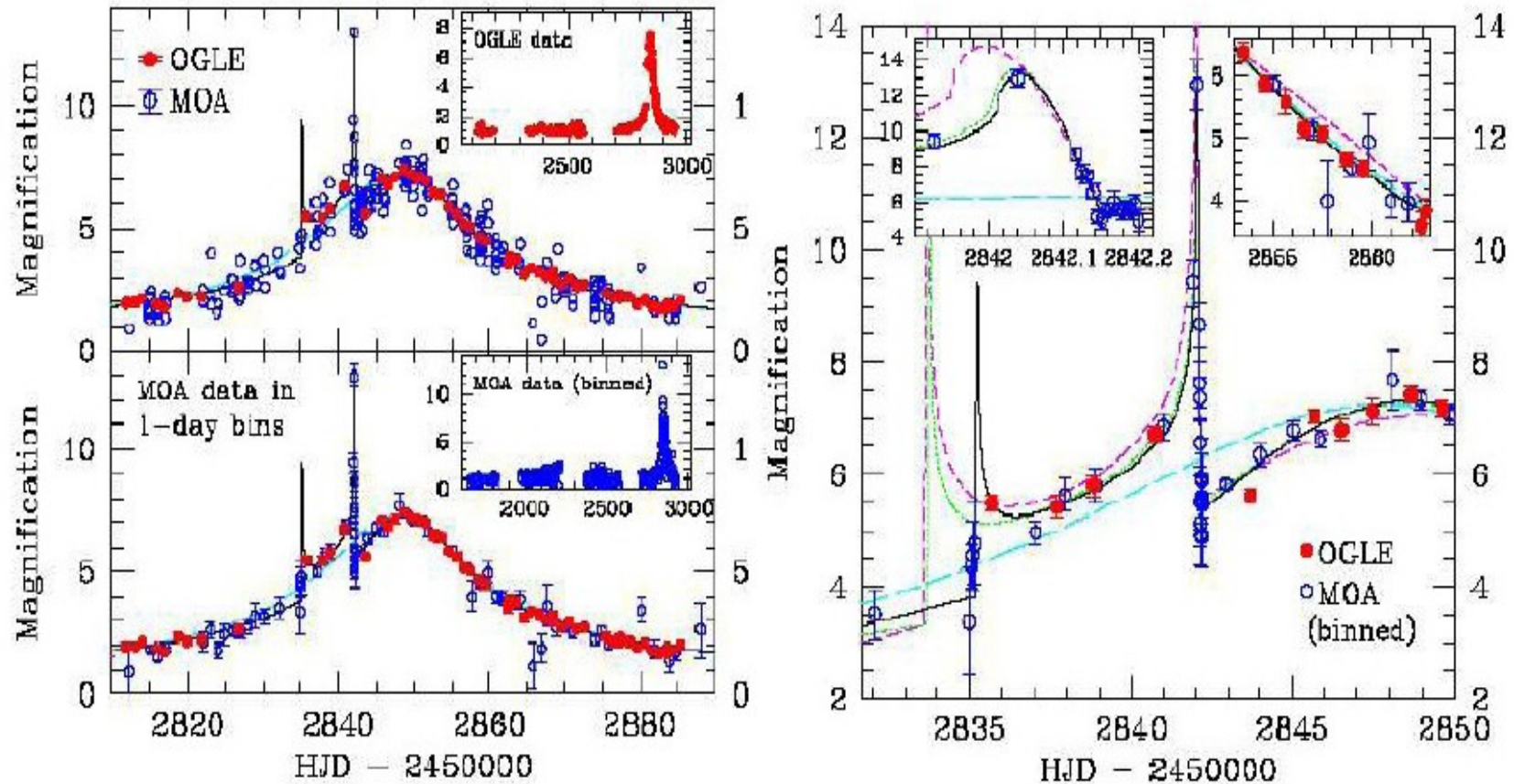


# CMD (Apparent Mags)



OB-03-235/MB-03-053: 5.5 kpc

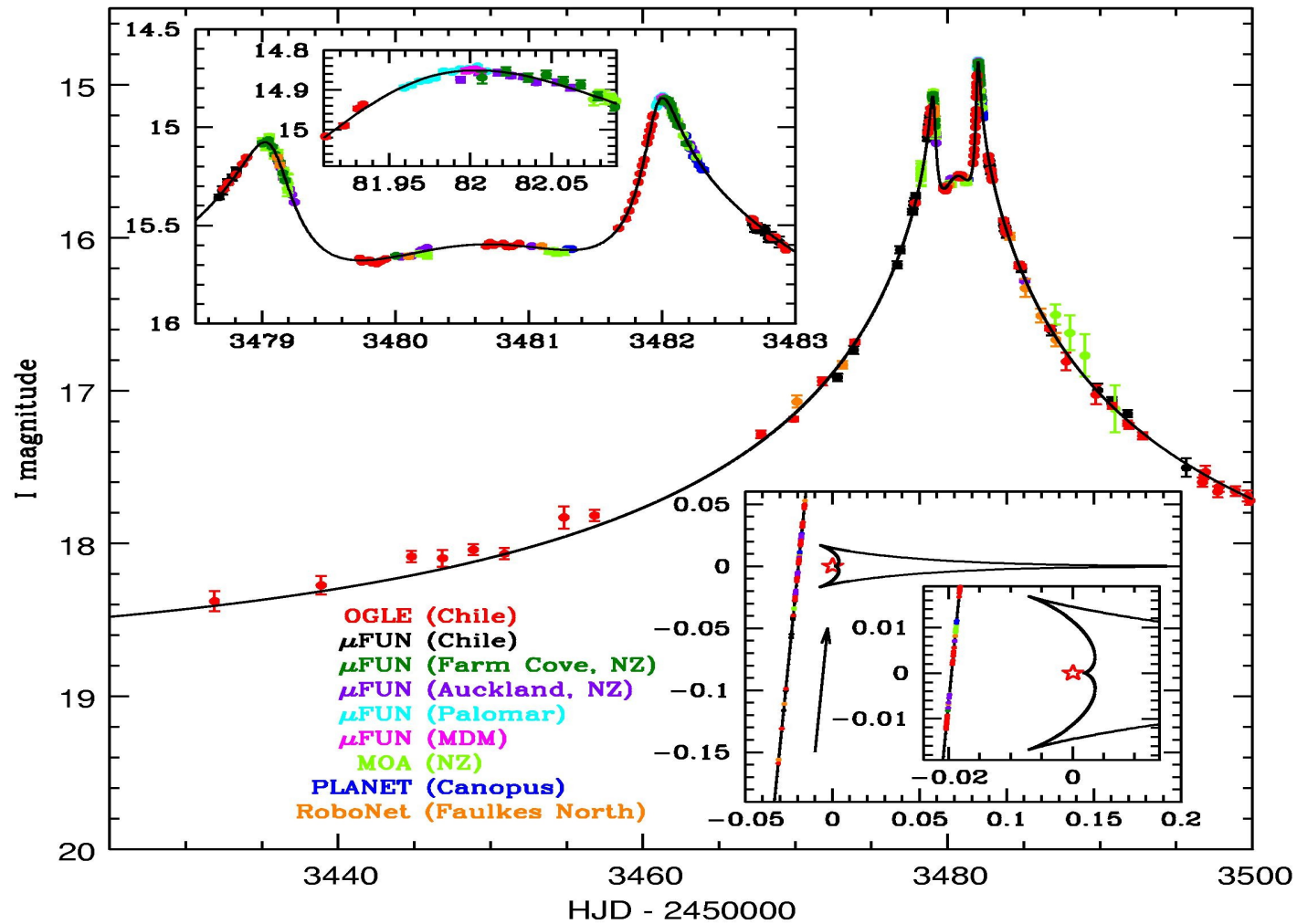
Finite Source + Centroid Motion



Bennett et al. 2006, ApJ, 647, L171

# OGLE-2005-BLG-071: 3.3 kpc

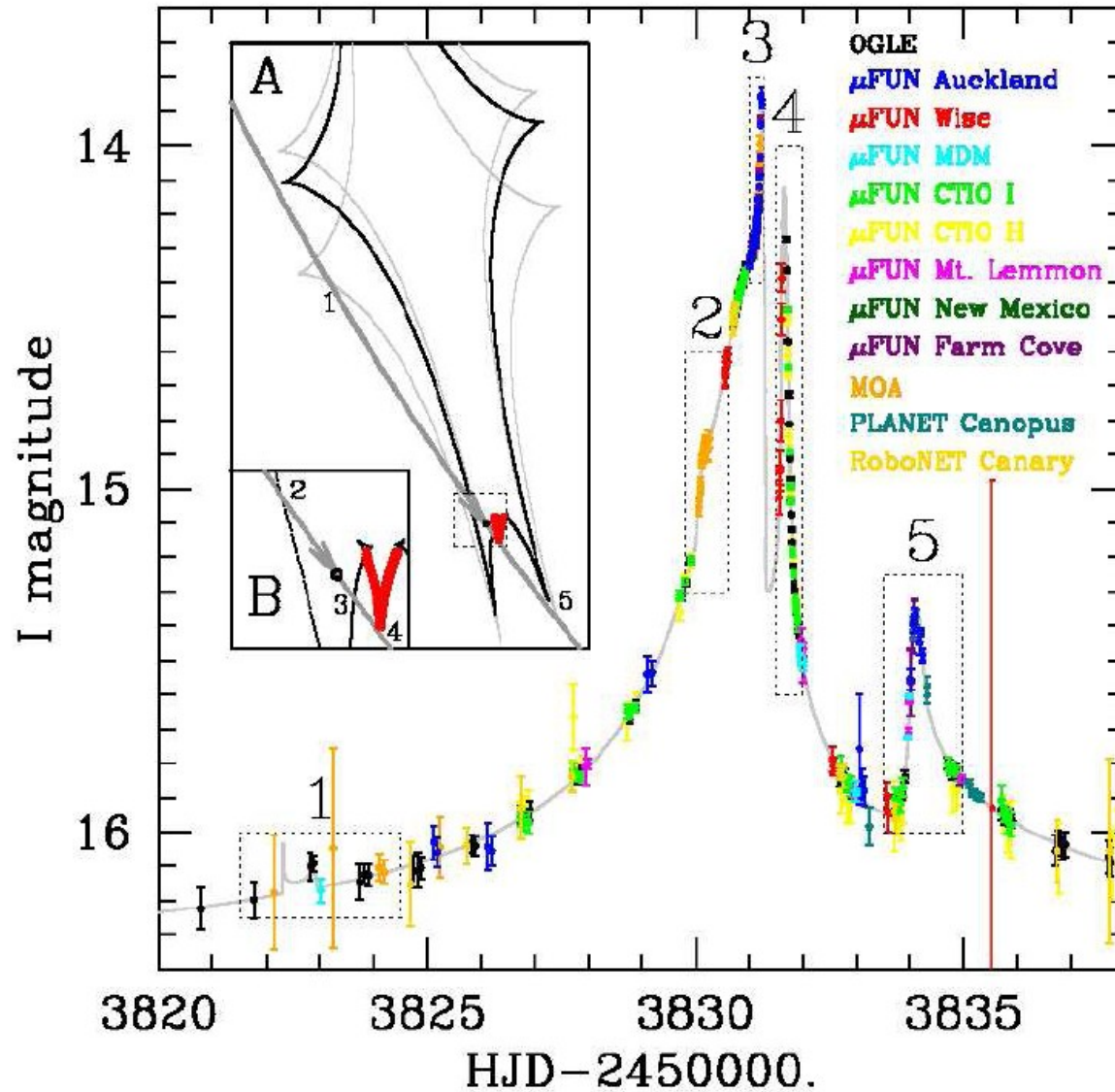
## Parallax + Finite Source + Centroid Motion



Dong et al. 2009, astro-ph/0804.1354

# OGLE-2006-BLG-109: 1.5 kpc

## Parallax + Finite Source + Blend

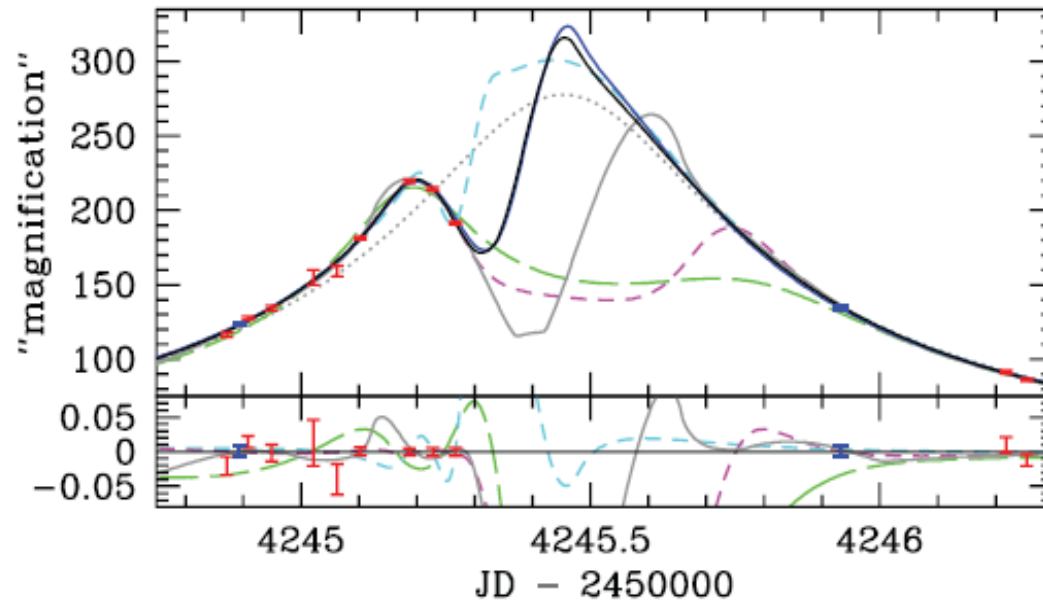
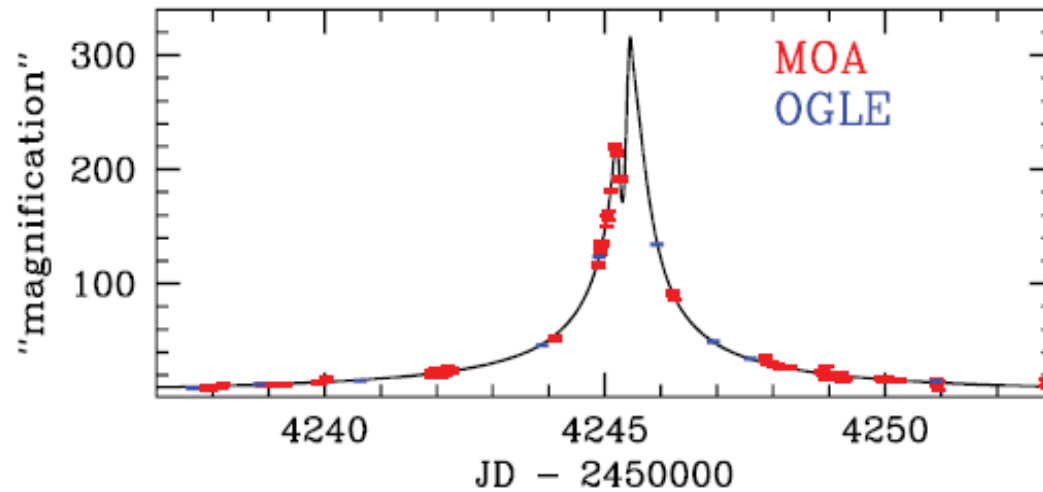


Gaudi et al. 2008, Science, 319, 927



# MOA-2007-BLG-192: 1.5 kpc

## Parallax + Finite Source



Bennett et al. 2008, ApJ, 684, 663

# Conclusions

- $\mu$ FUN has played a major role in the discovery of 9 extrasolar planets
  - 2005 (2), 2006 (2), 2007 (2), 2008 (3)
- Many planet distances are measured
  - Most in Disk
- Microlensing planets are cold
  - Great Probe Beyond the “Snow Line”