

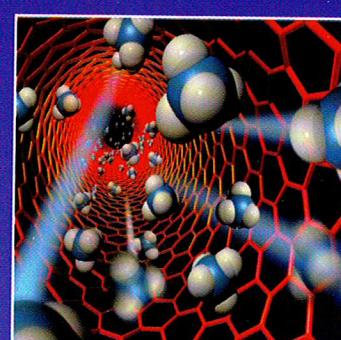
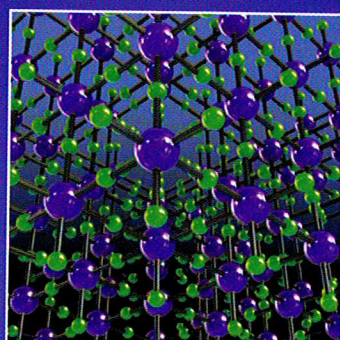
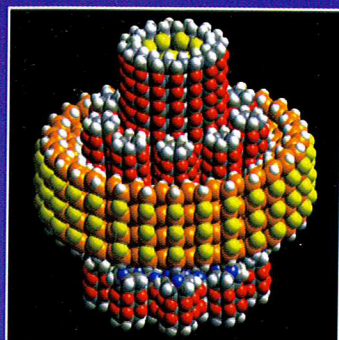
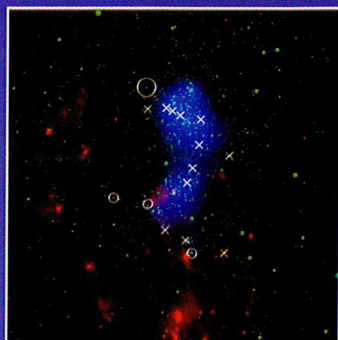


# CDAMOP 2011

December 14th–16th, 2011, New Delhi, India



Proceedings of  
3rd International Conference on Current Developments  
in Atomic, Molecular, Optical and Nano Physics with Applications



Organized by: Department of Physics & Astrophysics,  
University of Delhi, Delhi, India  
Co-sponsored by: LASTEC, DRDO, Delhi, India

# Photoionization, Photoexcitation and Astrophysical Opacities: The Iron Project

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Radiation transport through astrophysical plasmas is carried out by photon absorptions and emissions by the elements in the plasma and gives a measure of the plasma opacity. The dominate radiative atomic processes that govern opacity are photoionization and photoexcitations. The monochromatic opacity  $\kappa(\nu)$  at a photon frequency  $\nu$  in a plasma depends on the atomic parameters, oscillator strengths ( $f$ ) and photoionization cross sections ( $\sigma_{PI}$ ), of these processes. However, total monochromatic opacity is determined by the summed contributions of all possible radiative transitions from all ionization stages of all elements in the plasma source and hence requires large amount of atomic data. Accuracy of these data is crucial for accuracy of opacity. Accuracy in atomic parameters is also essential for various diagnostics of astrophysical plasmas. I will present results of photoionization and oscillator strengths from the recent developments in theoretical quantum mechanical calculations, under the international Iron Project and their impact on the astrophysical opacities and other astrophysical problems.

## Acknowledgement

Partially supported by DOE and NSF grants.

