Astronomy 161 -- Autumn 2007 relation of tectonism to internal heat In-Class Quiz 4 Study Guide different forces that shaped the different surfaces Origin & evolution of terrestrial planet atmospheres Probably started with similar compositions & evolved very differently General Solar System Role of the Greenhouse effect in atmosphere evolution _____ Names of the 8 planets Runaway Greenhouse effect Dwarf Planets (names of the 3 initial dwarf planets) Atmospheric Retention Order of planets in the Solar System The Habitable Zone Main types of planets & other bodies Jupiter & Saturn _____ Origin of the Solar System Gas Giants atmosphere composition _____ Clues to the formation in the present-day properties of the Solar System internal structure clues from the orbital motions internal heat clues from composition differences at different distances from the Sun atmospheric features Primordial Solar Nebula belts & zones Condensation of gases into solid grains as the Solar Nebula cools. winds The "frost line" Great Red Spot of Jupiter Formation of planetesimals from grains differences between Jupiter & Saturn Formation of the Terrestrial Planets magnetic field of Jupiter Formation of the Jovian Planets Formation of the Asteroids, Comets, and Kuiper Belt Objects Uranus & Neptune -----Ice Giants Terrestrial Planets Rocky cores and slushy icy mantles Extreme tilt and Extreme seasons of Uranus Mercury Lack of internal heat on Uranus and its effect on Uranus' weather Rotation Period locked in 3:2 tidal resonance with Sun Internal heat of Neptune and its effect on Neptune's weather Mercury's surface Impact craters Comparison of the Jovian Planets Caloris Impact Basin & jumbled terrain at the antipodes Mercury's atmosphere (such as it is) Differences in internal structure How did it get that way? Differences in amount of internal heat and its relation Mercury's interior to the different "weather" on each of the Jovian planets. large iron core and its supposed origin in a head-on impact Differences in magnetic fields weak magnetic field Venus Slow, retrograde rotation Venus' atmosphere composition pressure and temperature clouds How did it get that way? Runaway Greenhouse Effect Venus' surface radar mapping results rolling plains, valleys, & highlands extinct(?) volcanos & impact craters lack of plate tectonics upwelling & downwelling tectonism Contrast with the Earth Mars 2 moons: Phobos & Deimos Martian Atmosphere composition pressure & temperature dust storms How did it get that way? Martian Surface Features plains & cratered highlands volcanos (Olympus Mons & Tharsis Region) canyons & channels polar caps Water on Mars Comparison of the Terrestrial Planets Interiors & Surfaces (differences & similarities) cratering as a way to determine terrain ages