The final is Monday, March 12, 11:30 am - 1:18 pm, in the usual classroom.

**Review Sessions**

Ying Zu (TA) and I are available during office hours this week if you have questions about the course material or homework assignments, and we will hold Q&A review sessions on Friday and Saturday. *After* the final lecture on Friday, I will stay for another half hour and briefly review the new equations from this part of the course (on the back of this review guide) and the solutions to the homework assignments.

Both review sessions will be in McPherson Lab, room 4054.

*Friday, March 9, 5:00 pm - 6:15 pm:* I will hold a Q&A review session. It’s up to you to come in with questions, but you will probably learn from the answers to your classmates’ questions as well as your own.

*Saturday, March 10, 1:00 pm - 2:15 pm:* Ying Zu will hold a Q&A review session. We are going to request that the building be unlocked at that time, but if you do find that doors are locked, please go to the center door on the west side of the building, facing Smith Lab, and we will try to have someone there to let you in. You should try to get to the building slightly before 1 pm.

Attending one of the Q&A review sessions will almost certainly improve your performance on the final. The more reviewing you have done in advance, the more helpful it will be. You are welcome to attend both reviews if you like.

**Mechanics**

The exam will be mostly multiple choice, similar in style to the midterm, but longer in total number of questions, and with a few questions designed to take more time and thought to answer. There will also be an essay question, longer than the one on the midterm, worth roughly 20% of the exam score.

You may bring your page of handwritten notes from the midterm plus one new page (both sides) of handwritten notes, and a calculator.

*Please remember to complete your student evaluation (next sheet) and bring it with you to the final.*

**Topics and review advice**

The exam will focus on material from the second half of the course, beginning with General Relativity (Section 6 of the course notes) and running through Gravity Waves (Section 13). This material does draw on our treatment of Newtonian gravity and special relativity from the first half of the course, so in that sense the exam is necessarily “cumulative.”

The most useful way to review is (probably) to go carefully over the lecture notes and the solution sets to Homework 3 and 4, making up your handwritten notes as you go. You should also go over the solution to Part II of Homework 2, as the ideas from this part came back in the latter half of the course. Even if you have your own lecture notes, it is probably useful to look at my online notes alongside them as you review.

Pay particular attention as you review to how physicists and astronomers have built up the evidence for the existence of black holes over the last 100 years (almost), and the combination of theoretical arguments and astronomical observations that lead us to think that black holes exist and have the
properties predicted by General Relativity. Pay attention to the roles that radio astronomy, X-ray astronomy, and “traditional” (visible light) astronomy have played in this story, and the role that gravity wave astronomy might play in the future. Also pay attention to the empirical evidence for GR, especially those observations and experiments that indicate that it is a better description of gravity than Newton’s theory.

Names you should know: Newton, Einstein, Schwarzschild, Chandrasekhar, Zwicky, Oppenheimer, Wheeler, Hawking

Equations

In addition to the equations reviewed for the midterms, you should be familiar with the following equations — what quantities are represented by the variables, what principles the equations express, and how they can be used:

\[ E \approx \frac{GM^2}{R} \quad \text{(Homework 3, Part III)} \]
\[ L = \frac{1}{12} \dot{M}c^2 \]
\[ L_{\text{Edd}} = 3 \times 10^4 L_\odot \left( \frac{M}{M_\odot} \right) \]
\[ M = \frac{v^2r}{G} \]
\[ h \approx \frac{1}{5} \frac{R_{\text{Sch}}}{D} \]
\[ P \approx \frac{2\pi R_{\text{Sch}}}{c} \approx 10^{-3} \left( \frac{M}{10M_\odot} \right) \text{sec}. \]

You should remember that the Schwarzschild radius for one solar mass is \( 2GM_\odot/c^2 = 3 \text{km} \).
Astronomy 142: Student Evaluation

Mechanics

Your feedback on Astronomy 142 will be valuable in helping to shape the way the course is taught next time. Every time I teach the course, I make detailed notes for what I want to change the next time through, based partly on my own assessment but most of all on the feedback I get from student evaluations.

Thinking over the content of the course and evaluating its strong and weak points is also valuable as part of reviewing for the final exam. I therefore request that you set aside 20 minutes during your review for the final to complete this evaluation. Please bring it with you to the final, where I will have an envelope for collecting them. Your evaluation should be anonymous, and you may write it or type it as you prefer. I will not read any evaluations until after I have completed grading the course.

Please ALSO fill out the online SEI evaluations for this class, for which you should have received a link by email. The statistical information from SEI ratings is valuable both to me and to the Astronomy Department as a way of tracking improvement in the course or of flagging potential problems, and it provides our best gauge of overall student satisfaction. The SEI ratings thus complement the more specific feedback that you give me on this evaluation; I prefer to receive the more detailed feedback on paper rather than via the SEI online comments. SEI evaluations can be completed until midnight on Sunday, March 11.

Content

For your evaluation, please tell me anything you wish about what aspects of the course you liked and what aspects could be improved. I list some questions below to prompt your thinking; you do not need to answer all of them or restrict yourself to them. Please write or type your evaluation on a separate sheet.

1. What material did you find most valuable? What did we spend too much time on? What do you wish we had spent more time on?
2. What aspects of the lectures were most valuable? What aspects could be improved? Would a different balance between blackboard and Power Point be more effective?
3. Did you find the book a useful part of the course? Was the intersection between the book, the lectures, and the assignments adequate? Were the supplementary materials (course notes, other handouts) valuable/sufficient? Was the course web page valuable/sufficient?
4. Were the homework assignments useful for learning the course material and teaching you new things? What would make them better?
5. Any comments on the exams?

For your reference, I have attached on the back side of this evaluation the learning objectives for GEC courses in the Natural Sciences and the specific learning objectives that I identified for this course when I first proposed it. Feel free to comment on the degree to which the course did or did not achieve these objectives.

If you enjoyed the course, please encourage other students to take it! Professor Chris Kochanek will teach it this spring, and probably next year as well. If you are interested in taking another astronomy course, I am happy to give you advice. Options include special topic courses on Life In the Universe (A141) and Cosmology (A143) and the general lecture or (better) honors sections of Astronomy 161 or 162. I will be teaching A143 next fall (when its number will be A1143).
Learning Objectives for Natural Sciences GEC Courses

The general learning objectives for GEC courses in the Natural Sciences are:

1. To understand the basic principles and central facts of the physical and biological sciences, and their interrelationships.

2. To understand when, where, and how the most important principles and facts were discovered, thus understanding the key events in the history of science both as events in human history and as case studies in the methods of science.

3. To understand the interaction between science and technology.

4. To understand the social and philosophical implications of major scientific discoveries.

The specific learning objectives of *Astronomy 142: Black Holes* are:

- Qualitative physical understanding of Newton’s and Einstein’s theories of gravity, space, and time, the similarities and differences between them, and the senses in which Einstein’s theory has superseded Newton’s.

- Understanding of how Einstein’s theory leads to the prediction of black holes and of the properties it predicts black holes to have.

- Understanding of the interplay between gravity, pressure, and nuclear energy generation in governing the life cycle of stars, and of how and why the deaths of massive stars are expected to lead to the formation of black holes.

- Understanding of how astronomers discovered the first empirical evidence for black holes and of how they have set out to demonstrate the existence of black holes as conclusively as possible.

- Understanding of why supermassive black holes are thought to be the central engines of quasars, the most luminous objects in the cosmos, and of the observational methods that are used to study quasars and the dormant black holes they have left behind in the centers of galaxies.

- Understanding of the ways that advanced space missions currently under development might lead to deeper understanding of black holes, by measuring X-rays from gas falling towards the event horizon and by measuring gravity waves — propagating ripples in spacetime — produced by colliding black holes at the far edge of the universe.