

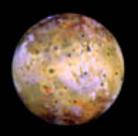
Ganymede 5262 km



Titan 5150 km



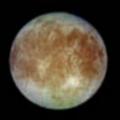
Callisto 4806 km



lo 3642 km



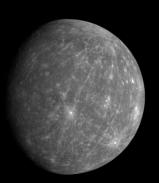
Moon 3476 km



Europa 3138 km



Triton 2706 km

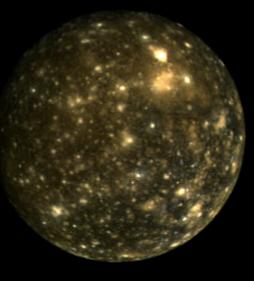


Mercury 4879 km

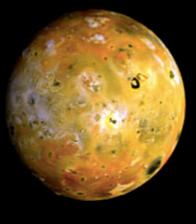
The Galilean Moons of Jupiter



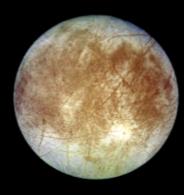
Ganymede (5262 km)



Callisto (4806 km)



Io (3642 km)



Europa (3130 km)



Moon (3474 km)

The Galilean Moons all orbit in the same direction around Jupiter.

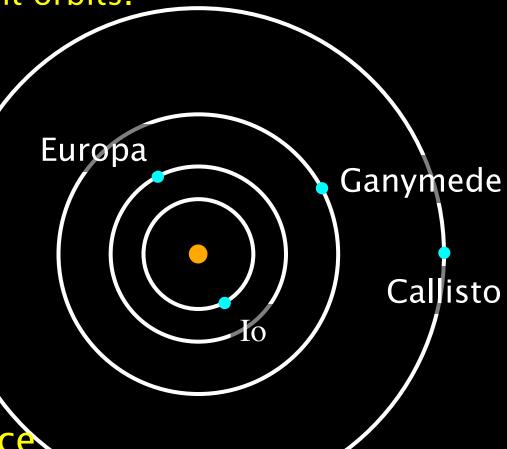
The inner 3 are on resonant orbits.

$$P_{lo} = 1.8^{d}$$

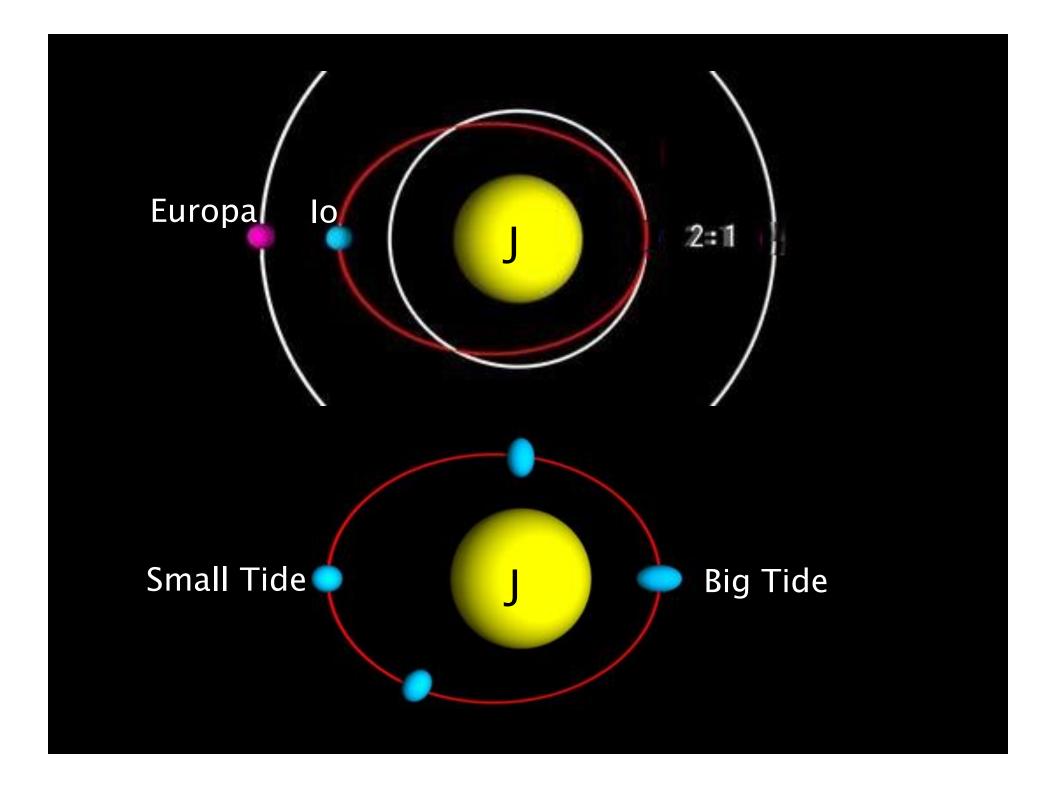
$$P_{Europa} = 3.6^d = 2 \times P_{lo}$$

$$P_{Ganymede} = 7.2^d = 4 \times P_0$$

$$P_{\text{Callisto}} = 16.7^{d}$$



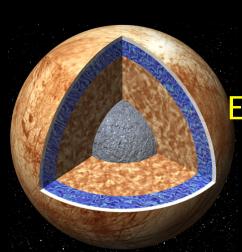
4:2:1 Laplace Resonance



Io and Europa

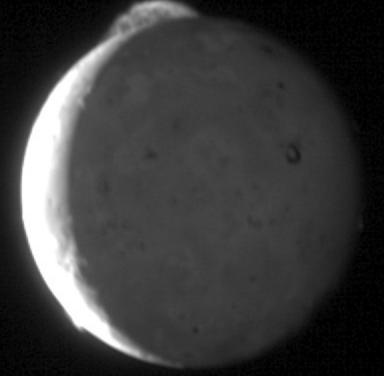
Mean densities of 3600 & 3000 kg m⁻³, respectively

lo: Rocky crust, molten mantle & many active volcanoes

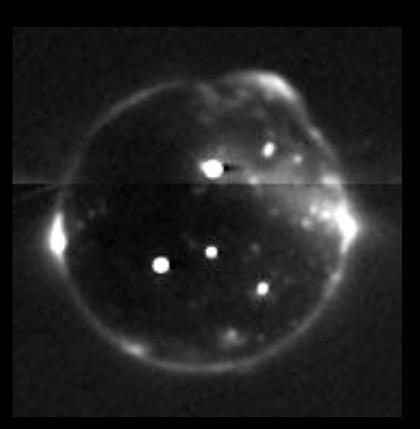


Europa: Icy lithosphere & rocky core. Likely has a deep-water ocean.

lo's active volcanoes



Tvashtar 2007 Feb 26 [New Horizons]



Io in eclipse, showing volcanic Hotspots 2007 Feb 27 [New Horizons]

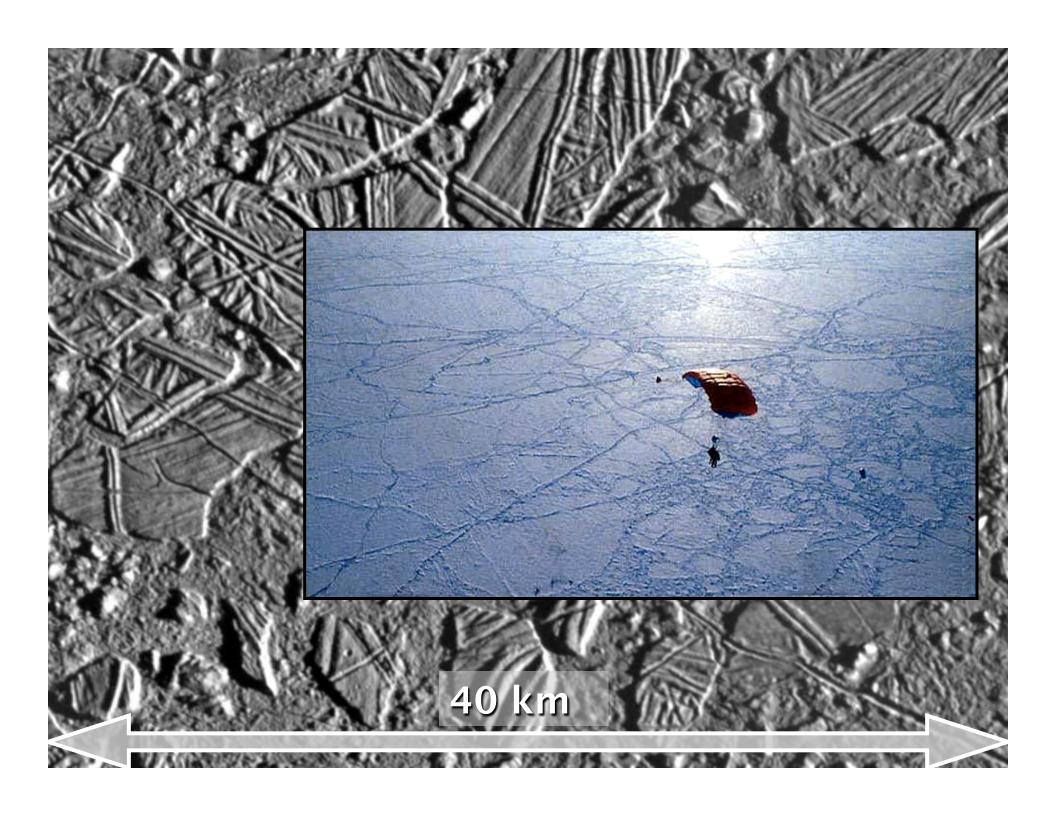
Europa has a smooth, young icy surface covering a large rocky core.

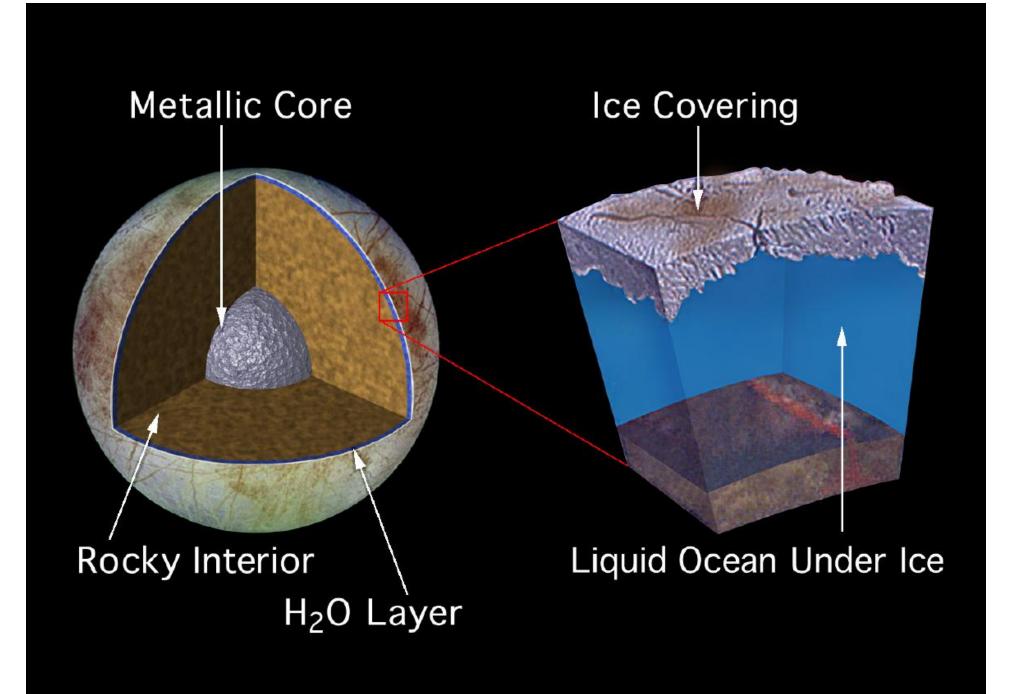
Composed of bright, shiny water ice.

Very few impact craters implies a young surface

Repaved by water geysering through cracks in the ice.

Ice surface is fractured into ice rafts and floes a few kilometers across

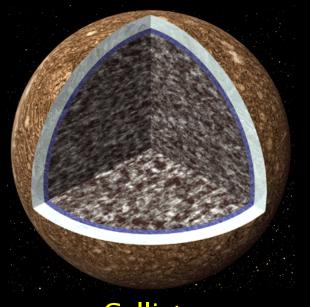




Ganymede & Callisto are mixed ice & rock, low-density moons.

Mean densities of ~1900 kg m⁻³

Deep ice mantles over rocky/icy cores.



Old, heavily cratered surfaces

Ganymede

They lack internal heat and are geologically inactive.

Callisto

Large Moons of Saturn

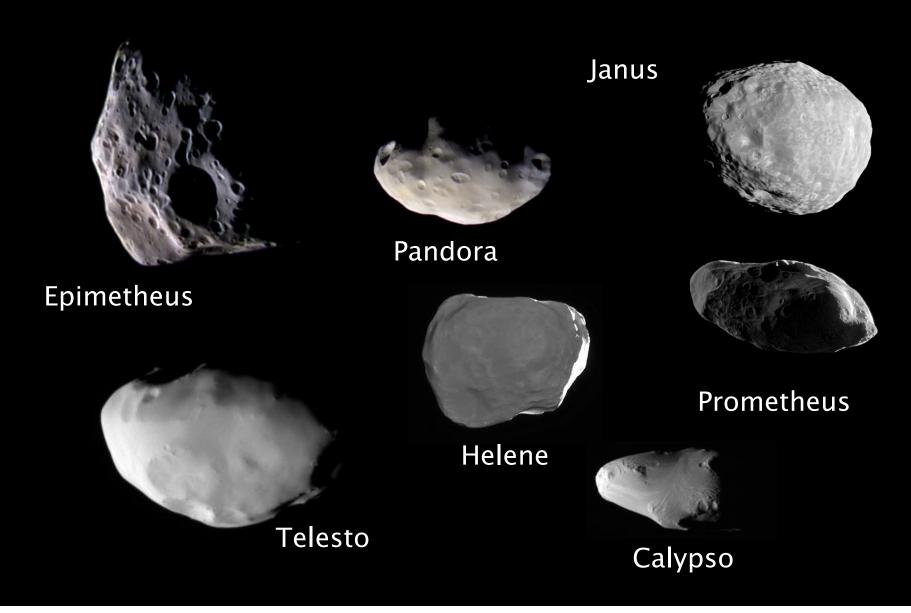
D > 200 km, mostly spherical

Titan



Tiny Irregular Moons of Saturn

D < 200 km



Enceladus is covered in fresh, clean ice.

Surface is lightly cratered, especially in the south.

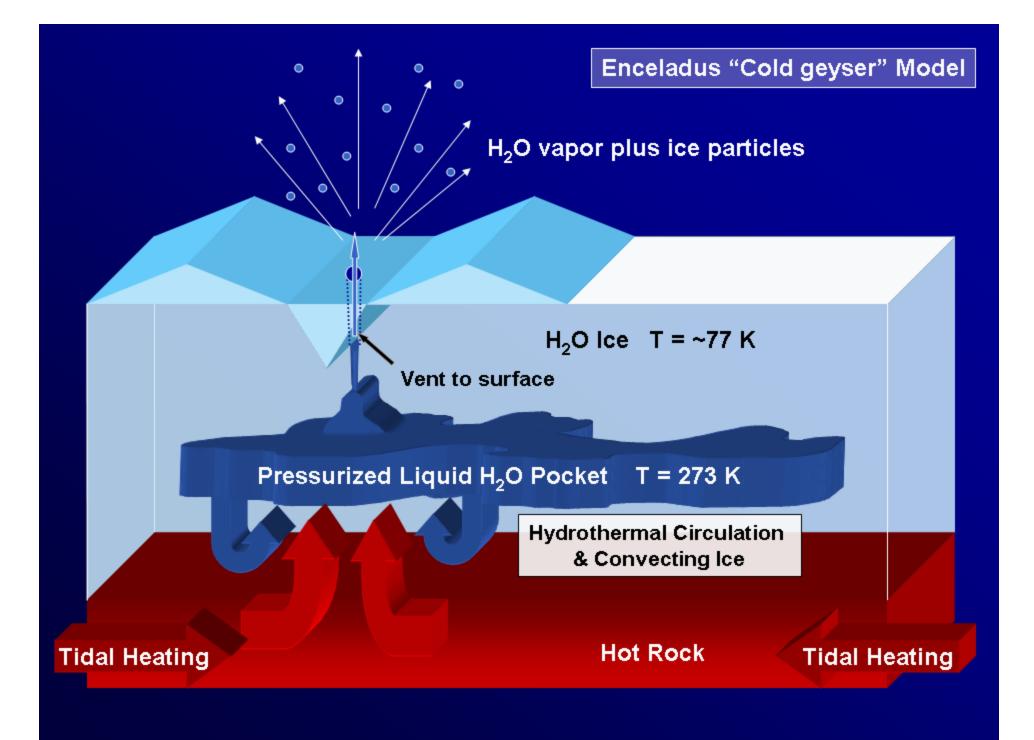
Tectonic features include scarps, grooves, and ridges, showing geologic activity.

A thin H₂O-vapor atmosphere & fresh surface ices fed by fountains at surface cracks.



The Fountains of Enceladus





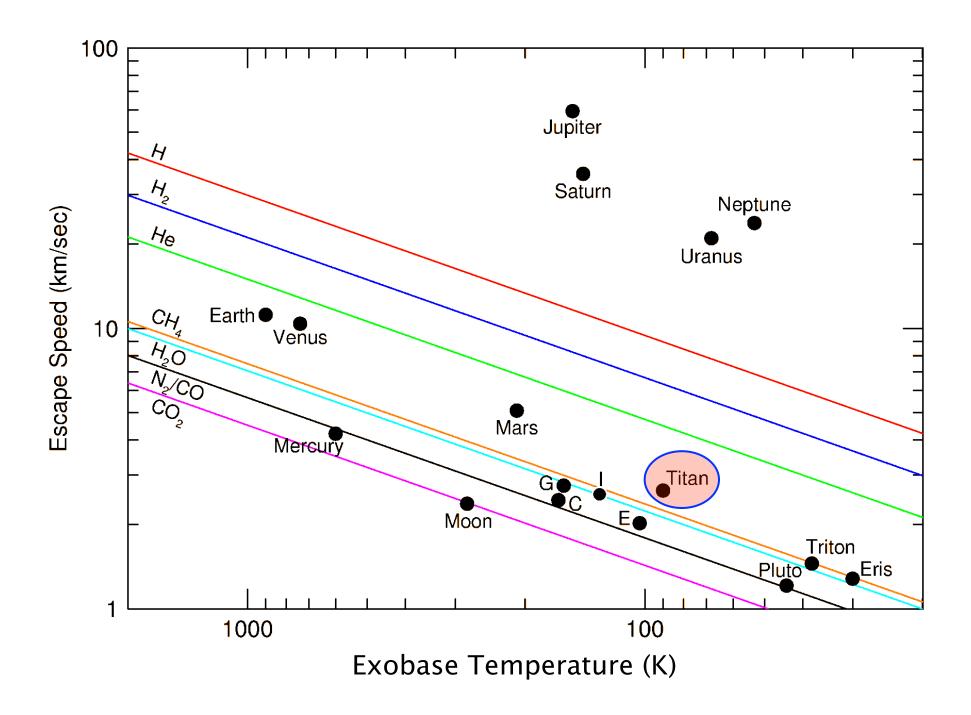
Titan

Radius: 2575 km

Density: ~1900 kg m⁻³ lcy mantle over a rocky core.

Cold enough to retain a heavy atmosphere of Nitrogen and Methane.

Pressure is high enough to have liquid methane on the surface.



Titan has a dense Nitrogen and Methane Atmosphere

Composition:

98% N₂ (nitrogen)

 $\sim 1.6\%$ CH₄ (methane)

Argon & hydrocarbons like Ethane

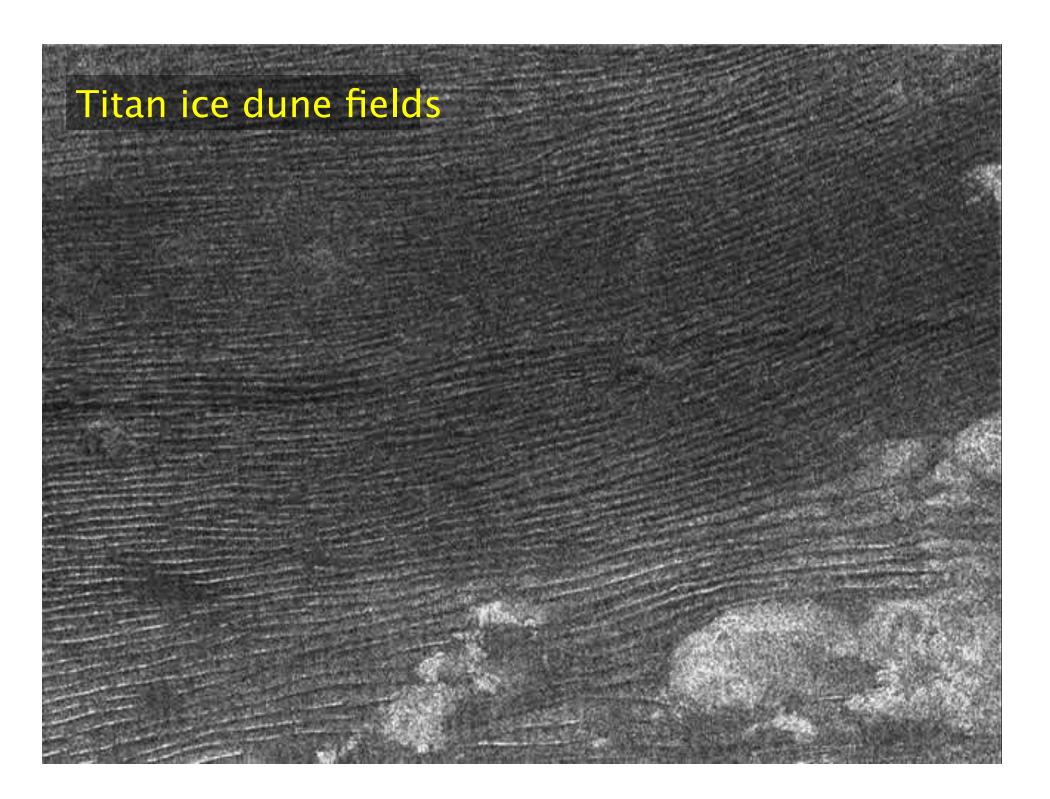
Cold and dense:

Temperature: 94 K (-290° F)

~1.6 Earth atmospheres pressure

Thick covering haze of brown photochemical aerosols (tholins)

Clouds of methane and ethane



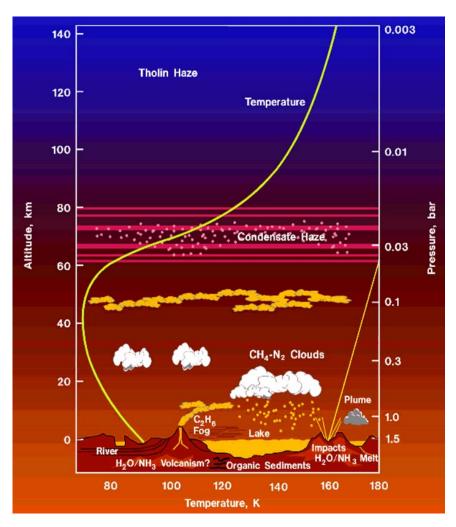
Methane (CH₄) plays the same role on Titan that water

does on the Earth.

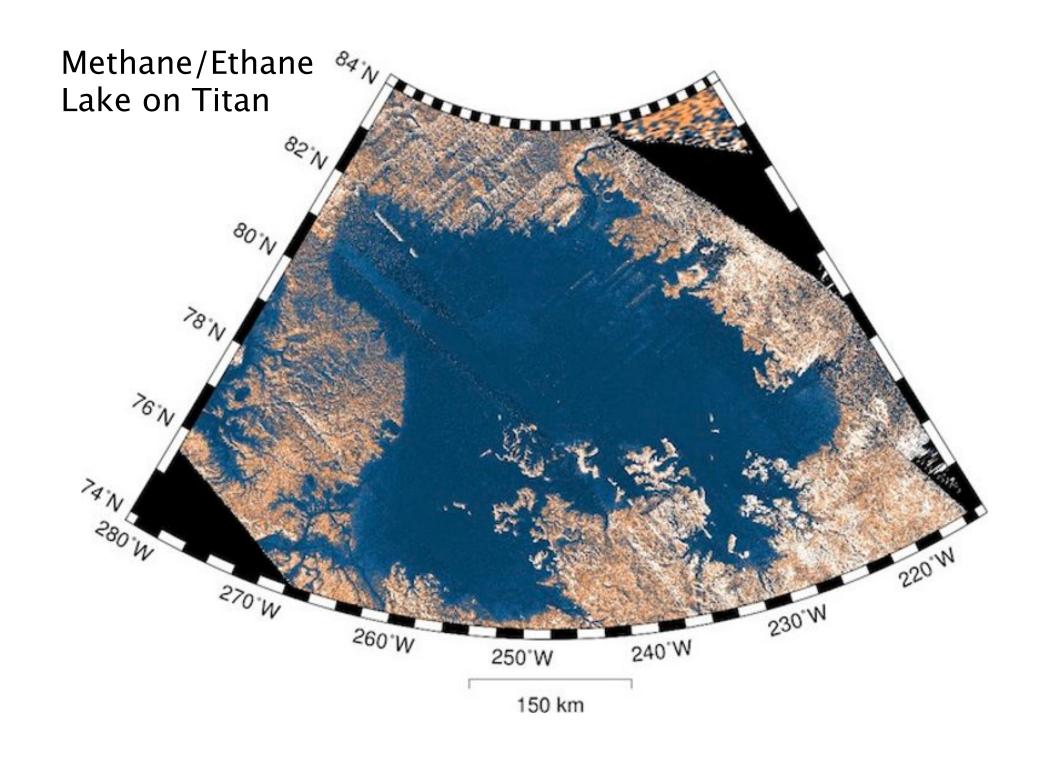
All three phases of methane exist at Titan's temperature & pressure

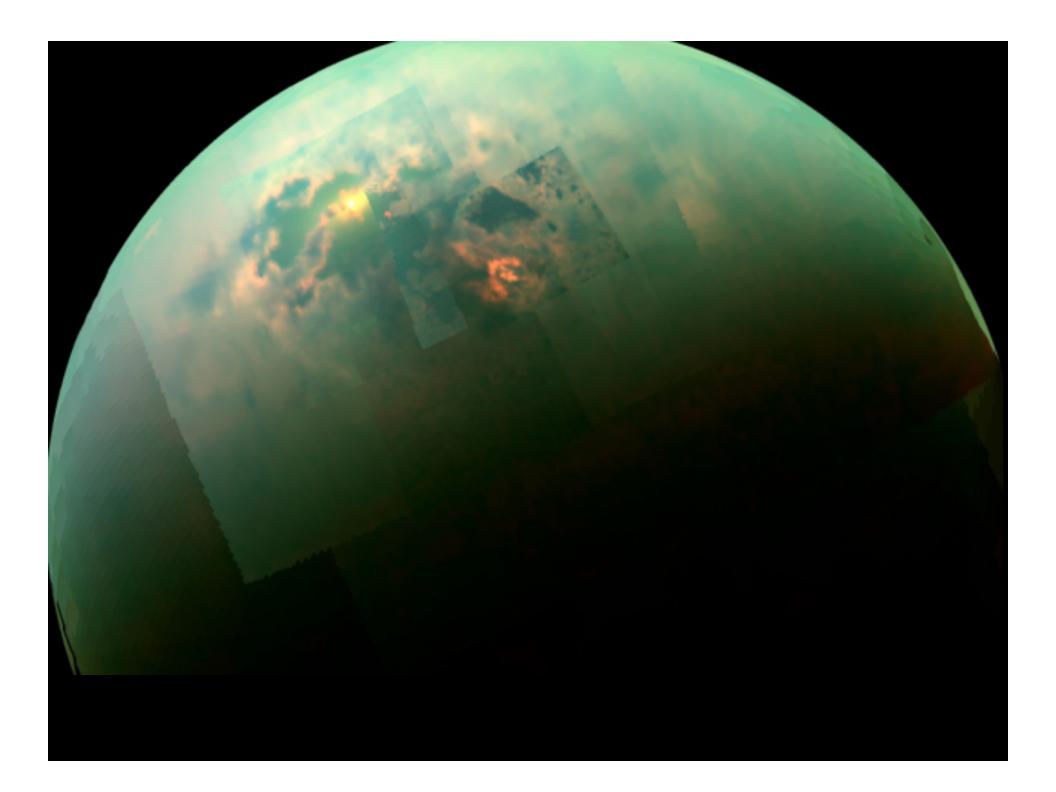
Atmospheric methane condenses into clouds that rain liquid methane.

Methane "Mud Flats" are water ice grains & liquid methane.



Liquid methane/ethane lakes found at the poles.



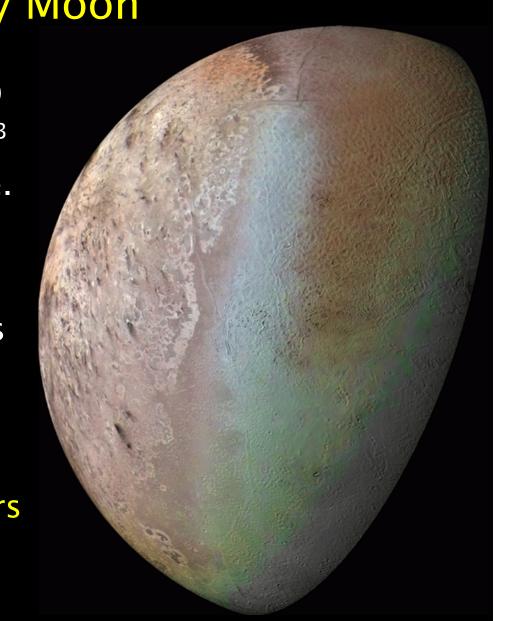


Triton: Neptune's Icy Moon

Diameter: 2710 km (21% R_E) Mean density: ~2050 kg m⁻³ Icy mantle over a rocky core.

Temperature 34 K (-398° F) N_2 , CH_4 , CO_2 , H_2O & CO ices Thin N_2 Atmosphere

Young surface with few craters



Smooth plains paved over by Cryovolcanic flows

N₂ Geysers:
Plumes of ices
& dark particles
Swept downwind,
making dark
streaks

Feeds Triton's thin N₂ atmosphere

