Jupiter's Dusty Rings

Faint, dusty rings:

Made of micron-sized dark dust particles. Mass is <10⁻¹² M_F

Sources:

Material knocked off small moons by meteoritic impacts Captured debris from comets

Jovian Ring System (schematic)





Saturn Ring System



67,000 km to 140,000 km (1.1 - 2.3 R_{Saturn}) ~30m thick

Water ice particles ~1 cm to ~3 m (pebbles to boulders) Main rings are inside Saturn's Roche Radius for icy bodies







Saturn's Rings backlit by the Sun (Cassini)





Thin Rings of Uranus Dark, narrow rings only a few km wide. Wide "epsilon ring" is only ~100 km wide.





Uranus' Rings back-lit by the Sun (Voyager 2)





Neptune's Rings



Galle Ring





Iceball-moon encounter:

Iceball on an inner (faster) orbit

Overtakes the moon and is deflected towards the moon



After the encounter, the ice ball falls back and oscillates around its original orbit

But as it passes back into the ring, it hits other iceballs, damping the oscillation

By the next encounter, it is back on its original orbit





Reference frame of the gap moon



Keeler Gap – cleared by small moon Daphnis



Thin rings can be gravitationally confined by a pair of "shepherd moons"

Outer moon



Shepherd moons of Saturn's F-ring

Pandora

Prometheus

http://photojournal.jpl.nasa.gov/animation/PIA07712

Close up of Prometheus and the F-Ring



Shepherd Moons of Uranus' Epsilon Ring

