

# Astronomy 350, Autumn 2002, Problem Set 5

Due Thursday, November 7 in class

## Problem 1

The thermal conductivity,  $Q$ , of a cylindrical bar of copper is proportional to its cross-sectional area,  $A = \pi r^2$ , where  $r$  = radius of the bar, and inversely proportional to its length,  $L$ :

- a) What is the fractional uncertainty of the derived thermal conductivity,  $\sigma_Q/Q$ , expressed in terms of the fractional uncertainties of the measured radius and length of the bar?
- b) Which quantity,  $L$  or  $r$ , should be measured with greater precision to better determine  $Q$ ? Please be quantitative (e.g., how much better)?

This kind of problem is an example of “working out the error budget” for a system.

## Problem 2

You have 10 images of a faint planetary nebula, and for each you have measured the apparent angular diameter,  $D$ , in arcseconds, and its uncertainty,  $\sigma_D$  as follows:

D (arcsec)	24.3	25.4	24.7	27.9	22.4	21.0	22.2	23.5	23.7	22.0
$\sigma_D$ (arcsec)	0.5	0.6	0.5	0.8	0.3	0.2	0.3	0.4	0.5	0.3

- a) Compute the weighted mean and its uncertainty from the data above.
- b) Compute the *unweighted* mean and its uncertainty. How does it compare to the weighted mean? What do you think is behind the difference?

## Problem 3

Fill in the right side of the table below with the error propagation formulae for common functional forms  $f(x,y)$ . Here,  $x$  and  $y$  are “observables” with uncertainties  $\sigma_x$  and  $\sigma_y$ , respectively, and  $a, b$ , etc. are constants with *no uncertainties*. In addition to computing  $\sigma_f$ , *if it is practical* also derive the fractional uncertainty ( $\sigma_f/f$ ) in terms of the fractional uncertainties of  $x$  and  $y$ .

Functional Form	Error Formula
$f = ax \pm by$	$\sigma_f =$
$f = \pm axy$	$\sigma_f =$
$f = \pm ax / y$	$\sigma_f =$
$f = \pm ax^{\pm b}$	$\sigma_f =$
$f = a \ln(\pm bx)$	$\sigma_f =$
$f = x \pm a \log_{10}(y)$	$\sigma_f =$

[Please make your own table, don't just fill in the blank example above, you will need more space.]

### Problem 4

The “activity” of a radioactive source is measured in terms of the number of decays observed per second. Over time, the activity decreases exponentially as

$$N(t) = N_0 e^{-t/\tau}$$

where  $N_0$  is the “initial activity” at time  $t=0$ , and  $\tau$  is the “lifetime” of the radioactive source. Suppose that for a particular experiment, both  $N_0$  and  $\tau$  are both known to a precision of 1%.

- a) Show that the uncertainty in the activity at a given time,  $\sigma_N(t)$ , is dominated by the uncertainty in the initial activity  $\sigma_{N_0}$  for early times  $t$ , and by the uncertainty in the lifetime  $\sigma_\tau$  at later times  $t$ .
- b) At what value of  $(t/\tau)$  do the uncertainties in  $N_0$  and  $\tau$  contribute equally to the uncertainty in the observed activity,  $\sigma_N$ ?

### Problem 5

The observed counts from an astronomical source are converted into a magnitude using a formula like the following:

$$m = m_0 - 2.5 \log_{10}(C)$$

where  $C$  is the observed counts, and  $m_0$  is the “Zero Point” of the photometric system.

- a) Suppose we know the zero point to perfect precision, and the observed uncertainty on the counts is  $\sigma_C$ . What is the uncertainty on the derived magnitude,  $\sigma_m$ ?
- b) Suppose that the zero-point has an uncertainty  $\sigma_{m_0}$ , and the uncertainty on the observed counts is still  $\sigma_C$ . What is  $\sigma_m$  now?