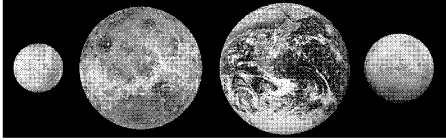


Wednesday, October 27  
Terrestrial Planets



**Announcement:** The question that was garbled on last Friday's quiz was also garbled in grading. The correct answer is "Each sequence of 3 base pairs in DNA molecules consists of a genetic **word** that codes for an **amino acid**." As a result of the erroneous grading, **one additional point** will be added to the score of everyone who took the quiz last Friday.

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Terrestrial Planets  
Key Concepts

- 1) Terrestrial planets all started out with thick atmospheres, but evolved differently.
- 2) **Small** terrestrial planets have old surfaces and cold interiors.
- 3) **Large** terrestrial planets have young surfaces and hot interiors.

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Terrestrial planets

**Large:**

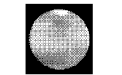
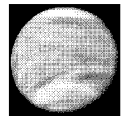
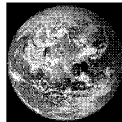
Earth (radius =  $1 R_E$ , mass =  $1 M_E$ )

Venus ( $0.95 R_E$ ,  $0.82 M_E$ )

**Small:**

Mars ( $0.53 R_E$ ,  $0.11 M_E$ )

Mercury ( $0.38 R_E$ ,  $0.055 M_E$ )



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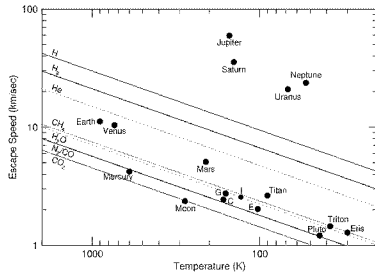
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The ability of a planet to retain an atmosphere depends on its mass and temperature.



Massive planets have a high escape speed; cold planets have slow-moving atmospheric molecules.

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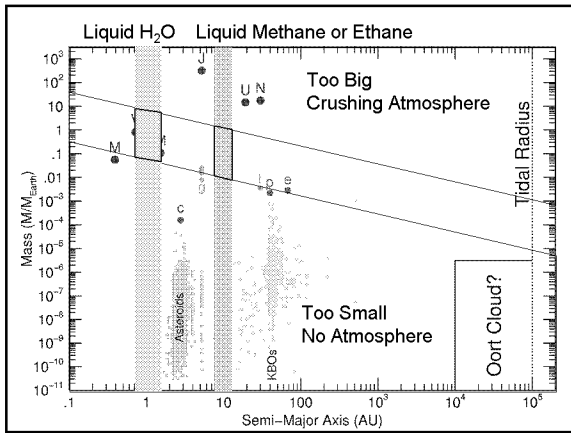
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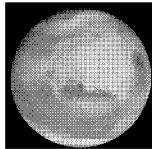
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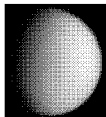
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Where should we look for life in the Solar System?

**Mars** may have had liquid water & a thicker atmosphere in the past.

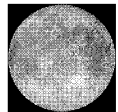


**Titan** (moon of Saturn) may have organic compounds in lakes of methane & ethane.



Wild cards:

**Europa** (moon of Jupiter) and **Enceladus** (moon of Saturn) may have liquid water under icy surfaces.




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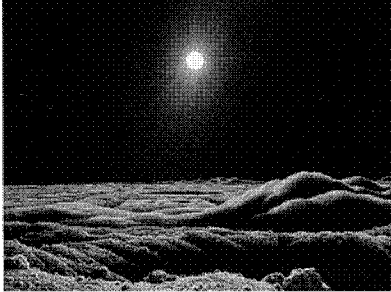
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**Mercury** is too hot for liquid water:  
it's too hot (and its gravity too weak) to  
have kept its primordial atmosphere.



Daytime highs = 425°C, nighttime lows = -175°C

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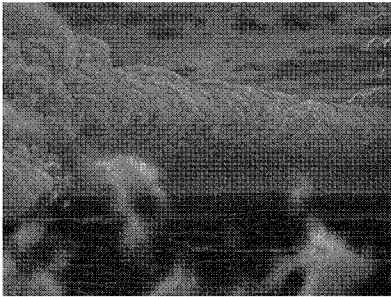
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**Venus** is too hot for liquid water:  
but its gravity is strong enough to  
retain its atmosphere.



Average temperature = 470°C

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Venus was once cooler, and may have had seas,  
but its temperature rose due to a  
**runaway greenhouse effect.**

Volcanoes belched out H<sub>2</sub>O, CO<sub>2</sub>, SO<sub>2</sub>  
(all greenhouse gases).

↓  
Air temperature rose.

↓  
H<sub>2</sub>O evaporated from seas;  
CO<sub>2</sub>, SO<sub>2</sub> released from seas & rocks.

↓  
Air temperature rose more.



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**Earth** has the right temperature for liquid water: its gravity is strong enough to retain its atmosphere.



**Average temperature = 15°C**

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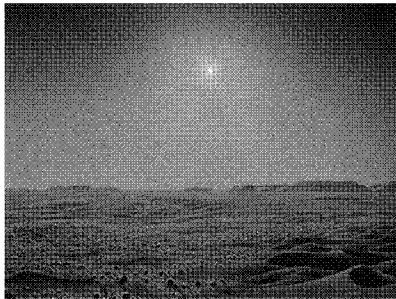
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**Mars** may have been warm enough for liquid water during its first Gyr, but its gravity is too weak to keep a thick atmosphere.



**Average temperature = -50°C**

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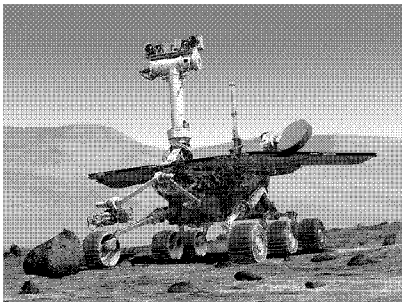
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Mars rovers have revealed evidence of past liquid water; as Mars cooled, the water froze.



**Mars has a cold, dry, thin CO<sub>2</sub> atmosphere today – but it might have been habitable in the past.**

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**Present-day** atmospheres are very diverse, despite similar **primordial** atmospheres.

	Earth	Venus	Mars
CO <sub>2</sub>	0.035%	96%	95%
N <sub>2</sub>	77%	3.5%	2.7%
H <sub>2</sub> O	1%	0.01%	0.007%
Ar	0.93%	0.007%	1.6%
O <sub>2</sub>	21%	trace	trace
	Habitable	Inhospitable today	

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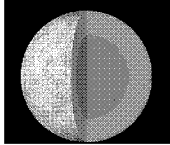
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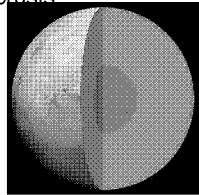
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The interiors of the **small** terrestrial planets cooled rapidly, and have solidified.

Mercury & Mars have thick, rigid crusts: no plate tectonics.



Mercury



Mars

Mercury has signs of ancient volcanic vents.  
Mars has large, extinct shield volcanoes.

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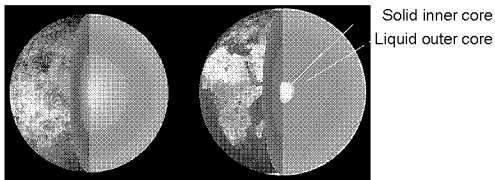
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The interiors of the **large** terrestrial planets cool more slowly, and are still partially liquid.

Venus & Earth are kept warm longer by energy release from radioactive decay.



Solid inner core  
Liquid outer core

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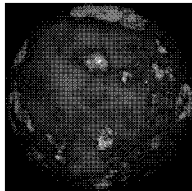
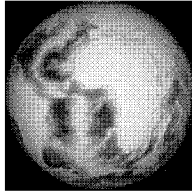
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The surfaces of the large terrestrial planets are **young**.

Average crust age:  
Earth: 200 Myr  
Venus: 500 Myr

**Earth:**  
plate tectonics,  
driving CO<sub>2</sub> cycle

**Venus:**  
lithosphere too soft  
to fracture into plates



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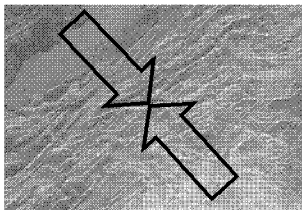
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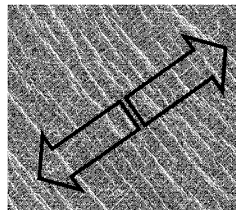
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There's evidence for **convection** in the mantle of Venus.



**Wrinkled mountains** are seen where the crust is compressed.



**Fractures** ("stretch marks") are seen where the crust is stretched.

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Tomorrow's Lecture:

**More about Mars**



This Week's Reading:

**Chapters 7 & 8**

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