Dealing with Rayleighs and Emission Measure: a worked example

In ISM studies, the standard unit of surface brightness for emission lines is the Rayleigh:

1 Rayleigh =
$$\frac{10^6}{4\pi}$$
 photons s⁻¹ cm⁻² sr⁻¹

The conversion between Rayleighs and surface brightness expressed in energy flux per square arcsec (e.g, erg s⁻¹ cm⁻² arcsec⁻²) depends on the wavelength of the emission line:

To convert, we need some useful numbers:

1 photon = hv or hc/ λ ergs = 1.98645×10⁻⁸/ $\lambda_{\text{Å}}$ ergs 1 steradian = 32400/ π^2 degrees² = 4.254517×10¹⁰ arcsec²

So

1 Rayleigh =
$$1.8704 \times 10^{-6}$$
 hv erg s⁻¹ cm⁻² arcsec⁻²

or

1 Rayleigh =
$$3.71546 \times 10^{-14} \lambda^{-1} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ arcsec}^{-2}$$

with λ in Angstroms.

For different recombination lines:

Thus

H
$$\alpha$$
: 1 Rayleigh = 5.661×10⁻¹⁸ erg s⁻¹ cm⁻² arcsec⁻²
H β : 1 Rayleigh = 7.643×10⁻¹⁸ erg s⁻¹ cm⁻² arcsec⁻²
Ly α : 1 Rayleigh = 3.056×10⁻¹⁷ erg s⁻¹ cm⁻² arcsec⁻²

The Emission Measure (EM), for pure recombination can be written:

$$EM = 2.75T_4^{0.9}I(H\alpha) \text{ cm}^{-6} \text{ pc}$$

= 7.84 $T_4^{0.9}I(H\beta) \text{ cm}^{-6} \text{ pc}$

for surface brightness, I(line) expressed in units of Rayleighs, or

$$EM = 4.858 \times 10^{17} T_4^{0.9} S(H\alpha) \text{ cm}^{-6} \text{ pc}$$
$$= 1.026 \times 10^{18} T_4^{0.9} S(H\beta) \text{ cm}^{-6} \text{ pc}$$

for surface brightness, S(line) expressed in units of erg s⁻¹ cm⁻² arcsec⁻².

For Ly α , the situation is a little more complicated. If we adopt the Case A ratio of Ly α /H β \approx 33, this implies

$$EM \approx 0.24T_4^{0.9}I(Ly\alpha) \text{ cm}^{-6} \text{ pc}$$