Light



Monday, October 5
Next Planetarium Shows: Tonight, Tue, Wed 7 pm

Universe contains electrically charged particles: protons (+) and electrons (-).

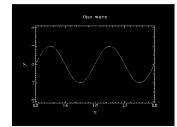
Charged particles are surrounded by electric fields and magnetic fields.

Fluctuations in those fields produce **electromagnetic waves**.

Visible light is a form of electromagnetic wave...

... but so are radio waves, microwaves, infrared light, ultraviolet light, X rays, and gamma rays.

Light is a wave.



Wave = a periodic fluctuation traveling through a medium.

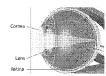
Ocean wave = fluctuation in height of water.



Sound wave = fluctuation in pressure.



Electromagnetic wave = fluctuation in electric and magnetic fields.



Describing a wave:

Wavelength (λ) = distance between wave crests.

Amplitude (a) = height of crests above troughs.

Frequency (f) = number of crests passing per second

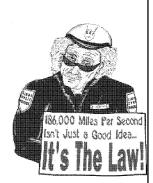
The speed of a wave equals wavelength times frequency.

$$c = \lambda \times f$$

(c for "celeritas", the Latin word for "speed")

| | wavelength | frequency | speed |
|---------------------------------|------------------------|---------------------------|-------|
| ocean wave | 100 meters | 0.1 /sec | |
| sound wave (middle C) | 1.2 m | 262 /sec | |
| light wave (red) | 6.6×10 ⁻⁷ m | 4.5×10 ¹⁴ /sec | |

The speed of light in a vacuum is **always** c = 300,000 km/sec (186,000 miles/sec).



| Light is made of particles . | |
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| | |
| Light shows some properties of particles, such as the photoelectric effect . | |
| UV light | |
| Particles of light, called photons , kick | - |
| electrons out of atoms. | |
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| The energy of a photon is related | |
| to the frequency of a wave. | |
| $E = h \times f$ | |
| | |
| E = energy of photon | - |
| f = frequency of light wave h = Planck's constant | |
| (a very small number indeed) | |
| | |
| | |
| | |
| | |
| Mayo or partialo? | |
| Wave or particle? | - |
| Both. | |
| Light has properties of both a wave and a | |
| stream of particles. Light follows the laws of | |
| quantum mechanics. | |
| 1 | |

Light forms a spectrum from short to long wavelength. 700 nm 600 nm 500 nm 400 nm

Visible light has wavelengths from 400 to 700 **nanometers**. [1 nanometer (nm) = 10⁻⁹ meters]

The **COMPLETE** spectrum of light

Gamma rays (λ < 0.01 nanometers)

 $X \; rays \; (0.01 \rightarrow 10 \; nm)$

Ultraviolet (10 → 400 nm)

Visible (400 \rightarrow 700 nm)

Infrared (700 nm \rightarrow 1 mm)

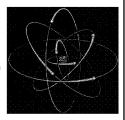
Microwave (1 \rightarrow 100 mm)

Radio (> 100 mm)

High energy photons

Low energy photons

Consider an atom: (highly schematic drawing)



A nucleus, consisting of protons and (usually) neutrons, is surrounded by a cloud of electrons.

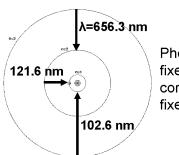
Hydrogen: one proton, one electron.



Behavior on subatomic scales is governed by **quantum mechanics**.

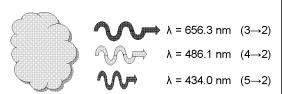
Rule: electrons can only exist in orbits of particular energy. (Small orbit = low energy, big orbit = high energy).

Electron falls from high- to low-energy orbit: energy is carried away by a photon.



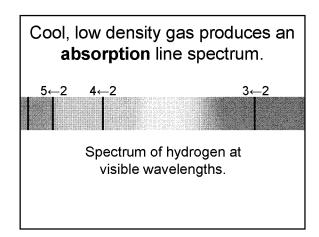
Photon has a fixed **energy**, corresponding to fixed **wavelength**.

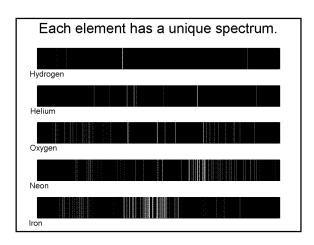
Consider a hot, low density glob of hydrogen gas.

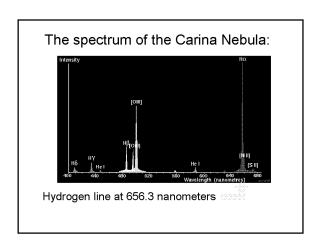


Light emitted **only** at wavelengths corresponding to energy jumps between electron orbits.

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|---|---|
| Hot, low density gas produces an emission line spectrum. | |
| - | |
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| Spectrum of hydrogen at visible wavelengths. | |
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| Carina Nebula: a cloud of hot, low density | |
| gas about 7000 light-years away. | - |
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| Its reddish color comes from the | |
| 656.3 nm emission line of hydrogen. | |
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| | |
| A cool, low density glob of hydrogen | |
| gas in front of a light source. | - |
| 100 → | |
| | |
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| | |
| Light shoothed only at way also after | |
| Light absorbed only at wavelengths corresponding to energy jumps | |
| between electron orbits. | |
| | |

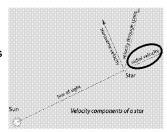






The **radial velocity** of an object is found from its **Doppler shift**.

Radial velocity = how fast an object is moving toward you or away from you.





Christian Doppler (1803-1853)

Doppler shift:

If a wave source moves toward you or away from you, the wavelength changes.

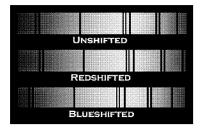
If a light source is moving **toward** you, wavelength is shorter (called "blueshift").

(should be "violetshift", more logically)



If a light source is moving **away** from you, wavelength is longer (called "redshift").

Doppler shifts are easily detected in emission or absorption line spectra.



Size of Doppler shift is proportional to radial velocity:

$$\frac{\Delta \lambda}{\lambda_0} = \frac{V}{c}$$

 $\Delta \lambda$ = observed wavelength shift = $\lambda - \lambda_0$

 λ_0 = wavelength if source isn't moving

V = radial velocity of moving source

c = speed of light = 300,000 km/sec

Wednesday's Lecture:

What is a star?

Reminders:

Have you read chapters 1 – 3?
Problem Set 1 is due Wednesday.
Planetarium shows Tonight, Tue, Wed.